

1       **The Sensitivity of Convection to Microphysics and Boundary Layer**

2               **Parameterizations in Hurricanes Harvey and Irma 2017**

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## ABSTRACT

10 Tropical cyclones (TCs) pose a significant threat to life and property, and exhibit many severe  
11 weather hazards as they make landfall, such as storm surge, strong winds, flooding rains, and  
12 tornadoes. TC convection is associated with nearly all of these hazards, which can extend hundreds  
13 of kilometers inland. Thus, understanding the characteristics and organization of convective cells is  
14 important to mitigating risk. Observational studies have noted that TC convection tends to organize  
15 downshear and that rotating thunderstorms tend to occur in the downshear-right quadrant of the TC.  
16 Modeling studies have also shown that convective cells tend to form upshear right and mature as the  
17 traverse cyclonically around the TC. Rotating thunderstorms in TCs are strongly influenced by the  
18 low-level helicity and convective available potential energy (CAPE), which have been highlighted  
19 in numerous modeling studies. The distribution and magnitude of low-level helicity and CAPE  
20 can be strongly influenced by microphysics and planetary boundary layer parameterizations in  
21 numerical weather prediction.

22 High-resolution Weather Research and Forecasting (WRF) simulations of hurricanes Harvey  
23 and Irma (2017) will investigate the role of microphysics and boundary layer parameterizations in  
24 determining the structure and distribution of rotating and non-rotating convection in TCs. Specif-  
25 ically, this project will examine how double- and single-moment microphysics parameterizations  
26 as well as local, non-local, and hybrid planetary boundary layer parameterizations impact the dis-  
27 tribution, structure, and longevity of convection. The high resolution (1 km) of these simulations  
28 will also allow for the investigation of whether boundaries at the TC or sub-TC scale influence con-  
29 vective organization. This study is unique in that it plans to investigate the interactions between  
30 microphysics and planetary boundary layer parameterizations on the development, evolution, and  
31 structure of both tropical cyclone convection and boundaries during landfall.



## 32 **1. Introduction**

33 Tropical cyclones (TCs) pose a significant threat to life and property for those living near the  
34 coast, exhibiting many different types of severe weather hazards as they make landfall, such as  
35 storm surge, strong winds, flooding rains, and tornadoes. Convection in tropical cyclones can con-  
36 tribute to a variety of these hazards. From 1995 to 2016, rotating convection in tropical cyclones  
37 directly resulted in 1296 confirmed tornadoes in the United States, accounting for 10–25% of all  
38 tornado activity in the coastal states from Louisiana to Maryland (Edwards 2012). Tropical cy-  
39 clone tornadoes also make up a large amount of the yearly tornado activity in Japan and China  
40 (Bai et al. 2019). Roughly 60% of landfalling tropical cyclones in the United States produce at  
41 least one tornado and the threat for such tornadoes can persist for up to five days after landfall  
42 (McCaul 1991). The risk for these tornadoes can extend 200–500 km from the tropical cyclone  
43 center to inland areas typically spared from strong winds and storm surge. The tornadoes associ-  
44 ated with tropical cyclones are typically weak with only 14% rated F/EF2 or higher (Schultz and  
45 Cecil 2009). Each tropical cyclone also has large variability in the amount of tornadoes reported.  
46 Some storms, such as Hurricane Ivan (2004), produce upwards of 118 tornado reports (Edwards  
47 2010), while others result in no tornado reports although sharing similar intensities and landfall  
48 locations. The weak and numerous tornadoes in tropical cyclones present a unique operational  
49 challenge to forecasters and decision makers as awareness may be relatively low compared to the  
50 other threats present in landfalling tropical cyclones (Weiss 1987; McCaul 1991).

51 The National Weather Service (NWS) performs service assessments to evaluate forecast per-  
52 formance following significant weather events such as hurricanes, floods, and impactful winter  
53 storms. Their assessment of Hurricane Irene (2011) discussed tornado warning false alarm rates  
54 of nearly 88%, which is above-average compared to all tornado warnings. It was also found that

55 the high false alarm rates of tropical cyclone tornado warnings damaged the credibility of the  
56 NWS (NWS 2012), taking away from other tropical cyclone risks. Martinaitis (2017) found a  
57 similar problem when looking at tropical cyclone landfalls from 2008 to 2013 in the United States  
58 that produced at least one confirmed tornado and in which at least 10 tornado warnings were is-  
59 sued. Martinaitis (2017) found that of the 1397 tornado warnings issued during the 12 tropical  
60 cyclones examined, only 198 tornado warnings verified, leading to an appalling false alarm rate  
61 of nearly 86%. In comparison, the national false alarm rate for tornado warnings in the United  
62 States have ranged from 80% in 1998 to 69% in 2016 (Fig. 1), which includes tropical cyclone  
63 tornado warnings. Brotzge et al. (2011) found that the false alarm rate for non-tropical cyclone tor-  
64 nado warnings from 2000 to 2004 was about 70%. Thus, tornado predication in tropical cyclones  
65 remains difficult.

66 According to the Storm Prediction Center (SPC), 2017 was the fourth most active year for tor-  
67 nado reports in tropical cyclones behind 2008 (third), 2005 (second), and 2004 (first). The two  
68 largest tornado producers of the 2017 tropical cyclone season were Hurricane Harvey and Hurri-  
69 cane Irma. Figures 2 and 3 show the locations of tornado reports during these two storms.

70 Current research, including this study, are focused on the use of high-resolution, convective-  
71 resolving models to study the formation, structure, and evolution, of rotating and non-rotating  
72 convection within tropical cyclones. A summary of previous observation and modeling studies  
73 that examined rotating convection in tropical cyclones, the effects of microphysics and planetary  
74 boundary layer parameterizations on numerical weather prediction, and boundaries within tropical  
75 cyclones will motivate this study.

## 76 **2. Literature review**

### 77 *a. Rotating convection in tropical cyclones*

78 Hurricane Danny (1985) was one of the first hurricane supercell environments to be studied  
79 comprehensively because of the 20 long-track supercells and 22 tornadoes reports it spawned (Mc-  
80 Caul 1987). McCaul (1987) noted that not only was veering of the low-level wind important, but  
81 so were dry air intrusions, which acted to increase convective instability. McCaul (1991) contin-  
82 ued this research by creating a climatology of buoyancy and shear in hurricane-spawned tornado  
83 environments using all available sounding data near reported tornado cases in the United States  
84 from 1948–1986. For the first time, it was documented that the distributions of buoyancy and  
85 shear in hurricanes had significant differences from quadrant to quadrant with respect to north, the  
86 direction of the large-scale vertical wind shear, and storm motion. The 0–3-km shear and helicity  
87 within the right-front quadrant was the most favorable for producing rotating convection and, in  
88 fact, these variables are very well correlated with the observed tornado frequency maximum in  
89 the right-front quadrant with respect to motion (McCaul 1991). In addition, climatological studies  
90 suggest that increased low-level shear is often associated with midlatitude tornado occurrences  
91 (Markowski et al. 2003).

92 Edwards (2012) reviewed the climatology, distributions, and environments of tropical cyclone  
93 tornadoes. In this review paper, the synoptic, tropical cyclone, and meso- $\beta$  scales were examined  
94 to summarize what influences tropical cyclone tornado and supercell potential on each scale. On  
95 the synoptic scale, the predominant driver of tropical cyclone convective (both rotating and non-  
96 rotating) development is the enhancement of vertical shear (McCaul 1991; Molinari and Vollaro  
97 2010). This increase in shear is generally attributed to tropical cyclone recurvature because of  
98 midlatitude westerlies and baroclinic boundaries. Consistent with Edwards (2012), Verbout et al.

99 (2007) found that tropical cyclones with relatively high tornado counts were accompanied by  
100 larger 500-hPa geopotential height anomalies and stronger height gradients.

101 Convection at the tropical cyclone scale is predominantly driven by the distributions of buoyancy  
102 and shear. Operational experience indicates that it is common for rotating convection to develop  
103 offshore and move inland. Some rotating convection weakens as it moves onto the more thermo-  
104 dynamically stable land as low-level (0–3-km) CAPE is about 35% less (Baker et al. 2009), while  
105 other convective cells increase mesocyclone intensity and undergo tornadogenesis due to the in-  
106 creased helicity from friction (Edwards 2012). On the meso- $\beta$  (convective) scale, tropical cyclone  
107 supercells have been observed to be smaller in vertical and horizontal extent compared to mid-  
108 latitude supercells (McCaul and Weisman 1996). Eastin and Link (2009) found in observations  
109 of Hurricane Ivan (2004) supercells were typically 5–7 km in diameter, compared to non-tropical  
110 cyclone supercells which are typically encompass a larger range of 3–12 km in diameter.

111 On the mesoscale, low-level, baroclinic, convergent boundaries and dry air intrusion can poten-  
112 tially influence the intensity and spatial distribution of tropical cyclone supercells (Edwards and  
113 Pietrycha 2006). Dry air ingested into the midlevels has a strong influence on convective struc-  
114 tures in tropical cyclones as it can substantially alter the vertical thermodynamic profile enhancing  
115 CAPE (McCaul 1987; Vescio et al. 1996; Curtis 2004). Dry slots can lead to the formation of  
116 baroclinic boundaries due to differential heating within the tropical cyclone envelope. Relatively  
117 cloud-free areas between tropical cyclone rainbands can support a few degrees Celsius of dia-  
118 batic surface heating (Card 2019). This surface heating can substantially magnify CAPE and yield  
119 baroclinic boundaries that may contribute to supercell maintenance (Edwards 2012). Edwards and  
120 Pietrycha (2006) argued that most landfalling tropical cyclones are not homogenized with equal  
121 tornado potential everywhere, and that boundaries and dry air intrusions may play a role in the  
122 clustering of tornadoes. Indeed, tropical cyclone tornado outbreak cases tend to have pronounced

123 relative humidity gradients from 700–500 hPa at the outer edge of the moist tropical cyclone en-  
124 velope (Curtis 2004)

125 Of tropical cyclones from 1948 to 2019, Hurricane Ivan (2004) holds the record for the number  
126 of tropical cyclone confirmed tornadoes at 118 (McCaul 1991; Schultz and Cecil 2009). Baker  
127 et al. (2009) looked at the environmental ingredients for the development of supercells and torna-  
128 does in Hurricane Ivan via airborne and land-based observations. The azimuthal location of the  
129 tornadoes in Hurricane Ivan could be explained by significant 0–1-km shear ( $7.4 \frac{m}{s}$ ) and low lifting  
130 condensation level (LCL) heights (415 m) in the right-front quadrant with respect to storm mo-  
131 tion. Motivated by an apparent increase in individual convective cell rotation as convection made  
132 landfall, Baker et al. (2009) further investigated the differences in the convective environments be-  
133 tween the sea and land. They found that the land soundings had very similar total-column CAPE  
134 to the sea soundings; however, the low-level (0–3-km) CAPE was 35% less over the land. McCaul  
135 and Weisman (1996, 2001) suggested that updraft strength and vorticity were both enhanced when  
136 buoyancy is concentrated in the low-levels, suggesting convection is more likely to form over the  
137 ocean and move onshore. The other appreciable difference between the land and sea environments  
138 in Baker et al. (2009) was that the 0–1-km storm relative helicity (SRH) was 50% greater over  
139 land, due to frictional effects.

140 Although not observed in Hurricane Ivan, some researchers have suggested that changes in  
141 surface wind speeds as large as  $8\text{--}10 \frac{m}{s}$  could occur across horizontal distances of 10 km at land–  
142 ocean interfaces (Powell and Houston 1998). Gentry (1983) showed that there is an increase  
143 in low-level helicity because of the increase in friction between the land–sea interface acting to  
144 enhance low-level vertical shear. As a result, individual convective cells making landfall tend to  
145 increase updraft rotation and intensity due to the enhanced low-level shear (Baker et al. 2009).

146 Eastin and Link (2009) used the same collection of airborne and land-based observations as  
147 Baker et al. (2009), and concluded that the offshore environment was conducive for supercell  
148 formation. In the examination of the individual rotating convective cells, mesocyclonic updrafts  
149 extended from the boundary layer up to 6–8 km and were 5–7 km in diameter. The production of  
150 the updraft likely results from a combination of convergence, thermal instability, and perturbation  
151 pressure gradients, which help to produce mesocyclones by tilting and stretching environmental  
152 vorticity (Eastin and Link 2009).

153 These observational studies of Hurricane Ivan (2004) led to high-resolution, real-data simula-  
154 tions to document the structure of potentially tornadic supercells embedded within tropical cy-  
155 clone rainbands. Carroll-Smith et al. (2019) produced one such simulation at 3- and 1-km grid  
156 spacing. In an attempt to verify the tropical cyclone tornadoes associated with Hurricane Ivan  
157 (2004), percentile values of maximum updraft helicity and simulated radar reflectivity were used  
158 to identify tropical cyclone tornado surrogates and compare those surrogates to observed tornado  
159 reports. The surrogates with the 99.9<sup>th</sup> (99.95<sup>th</sup>) percentile of maximum updraft helicity in the  
160 3-km (1-km) domain provided the most favorable results capturing the distribution of tropical cy-  
161 clone tornadoes compared to observations. These high updraft helicity percentiles suggest that  
162 supercells with strong mesocyclones are more likely to produce tornado reports in tropical cy-  
163 clones. In this modeling study of Hurricane Ivan, updraft helicity and simulated radar reflectivity  
164 were used successfully as tropical cyclone tornado surrogates (Carroll-Smith et al. 2019).

165 Card (2019) used a similar analysis technique to Carroll-Smith et al. (2019) to diagnose rotating  
166 convection in hurricanes Harvey and Irma (2017) using the the National Center for Atmospheric  
167 Research (NCAR) 10 member ensemble. In Card (2019), the number of identified rotating storms  
168 outnumbered the identified non-rotating storms by a factor of 2–3 in both Harvey and Irma (2017).  
169 With respect to storm motion and north the distributions of rotating and non-rotating convection

170 is very similar (Figs. 4 and 5). There is a strong relationship between shear and storm motion  
171 near the US coasts because tropical cyclones are typically recurving. As shown in Corbosiero and  
172 Molinari (2003), shear is the dominate factor in the distribution of convection in tropical cyclones.  
173 Most of the rotating storms occur directly downshear, while most of the non-rotating storms occur  
174 upshear-right in both the NCAR ensemble and in observations (Figs. 4 and 5).

175 In summary, the common environmental characteristics of tornadic rotating convection in trop-  
176 ical cyclones are: 1) high 0–3-km storm relative helicity (SRH), 2) high 0–3-km CAPE, 3) low  
177 lifting condensation level (LCL) heights, 4) relatively dry air at midlevels, and 5) low-level bound-  
178 aries (such as a convergence axis or baroclinic zone) (Novlan and Gray 1974; McCaul 1991; Curtis  
179 2004; Edwards and Pietrycha 2006; Eastin and Link 2009).

#### 180 *b. Microphysics sensitivity in numerical weather prediction*

181 Microphysical schemes parameterize many different small scale processes dealing with precipi-  
182 tation. Microphysics schemes track a number of different species of hydrometers, phase changes,  
183 and information about the mass, number, and size of the hydrometers. Single moment micro-  
184 physics schemes predict the total mass concentration of hydrometers, while double moment mi-  
185 crophysics schemes often include a prediction of the total number concentration for some species  
186 of hydrometer in addition to mass concentrations. Both the WRF single moment 6-class (WSM6)  
187 and the WRF double moment 6-class (WDM6) schemes track the mixing ratios of six different  
188 hydrometer species (water vapor, clouds, ice, snow, rain, and graupel) (Hong and Lim 2006; Lim  
189 and Hong 2010). WDM6 is double moment for warm rain processes, meaning it additionally pro-  
190 vides prognostic number concentrations of cloud and rain water, as well as cloud condensation  
191 nuclei (CCN) (Lim and Hong 2010). The predicted CCN number concentration in the WDM6  
192 microphysics scheme adds a level of complexity to traditional bulk microphysics schemes through

193 explicit CCN–cloud drop concentration feedbacks. An example of this is assuming evaporation  
194 of cloud drops returns the corresponding CCN particles to the total CCN count. The warm rain  
195 source and sink terms are the same for WSM6 and WDM6; however, WDM6 uses auto-conversion  
196 and accretion based on Cohard and Pinty (2000). Many of the microphysical processes in WDM6  
197 use the same formulas as WSM6, although they work differently due to the predicted number  
198 concentrations of cloud water and rain, which can indirectly influence ice processes (Hong et al.  
199 2010).

200 Microphysical parameterizations can have a large impact on the vertical structure and develop-  
201 ment of individual convective cells. It is also well documented in the literature that microphysical  
202 parameterization have an impact on tropical cyclone intensity and track (Willoughby et al. 1984;  
203 Lord and Lord 1988; Zhu and Zhang 2006; Fovell and Su 2007; Li and Pu 2008; Fovell et al.  
204 2009; Tao et al. 2011; Fovell et al. 2016). The Korean Meteorological Administration (KMA) has  
205 used both WSM6 and WDM6 microphysics schemes operationally. The KMA has shown no dis-  
206 tinct discrepancies in predicted precipitation between these two schemes, but WDM6 has shown  
207 superior predictive skill in a variety of weather conditions (Hong et al. 2010).

208 The results of Hong et al. (2010) also noted that the WDM6 scheme tends to suppress spurious  
209 light precipitation over oceans. Since WDM6 tends to suppress spurious light precipitation that  
210 occurs over the ocean, it may affect the degree of clearing seen between rainbands, which can  
211 impact the amount of baroclinic forcing and CAPE on the radially inward and outward sides  
212 of the rainband convection (Hong et al. 2010; Yussouf et al. 2013). Additional clearing would  
213 favor more convectively active rainbands with the possibility of more rotating and non-rotating  
214 convective cells.



215 Hong et al. (2010) also reported that the WDM6 tends to propagate squall-lines too quickly.  
216 Distant convective rainbands can sometimes take on squall line-like properties (Houze 2010) and  
217 may propagate outward radially too quickly when using the WDM6 microphysics scheme.

218 Microphysical parameterizations are one important source of error in storm-scale modeling at  
219 high resolution. For example, Yussouf et al. (2013) examined a tornadic supercell from 8 May  
220 2003 in Oklahoma City using single and double moment microphysics schemes. The double mo-  
221 ment scheme supported a better distribution of the reflectivity in the forward flank region of the  
222 simulated supercells than the single moment scheme. Putnam et al. (2017) used 4-km Storm-Scale  
223 Ensemble Forecasts (SSEF) to simulate polarimetric radar variables and compared those with ob-  
224 servations, specifically looking at the simulation hydrometeor types and particle size distributions.  
225 Two particular cases, both from 20 May 2013, were examined, the first being a mesoscale con-  
226 vective system and the second being a supercell thunderstorm. Putnam et al. (2017) found that  
227 WSM6 had poor coverage of stratiform precipitation. Despite being double-moment for warm  
228 rain processes, WDM6 had a similar relationship to WSM6 with respects to simulated reflectivity  
229 and differential reflectivity (Putnam et al. 2017). All of the double-moment microphysics schemes  
230 tested in Putnam et al. (2017) exhibited incorrect differential reflectivity maxima associated with  
231 isolated, weak convection on the back side of the convective lines where large raindrops would not  
232 be expected. In both the mesoscale convective system and in the supercell cases, WSM6 produced  
233 mainly rain while WDM6 produced mainly rain and graupel. In the supercell case, the WDM6  
234 produced much less reflectivity than the other simulations. They also found that both WSM6 and  
235 WDM6 have a bias toward small raindrops and graupel (Putnam et al. 2017).

236 Numerous studies have highlighted the effects of microphysical parameterizations on tropical  
237 cyclones. Fovell et al. (2009) demonstrated that varying microphysics can result in different wind  
238 profiles 100–300 km from the storms center, which directly influences the track. Track variations

239 with respect to different microphysics schemes disappear when hydrometeors can no longer inter-  
240 act with longwave and shortwave radiation (Fovell et al. 2010). Tropical cyclone intensity is also  
241 influenced by microphysics. In general, the exclusion or reduction in graupel in the cloud results in  
242 an increase in intensity and tangential winds in tropical cyclones (McFarquhar et al. 2006; Fovell  
243 et al. 2009). Lastly, WDM6 and WSM6 will likely have different cloud, rain, and ice concentra-  
244 tions. Microphysics parameterization produces different concentrations, types, and distributions  
245 of hydrometeors which has an impact on storm dynamics and thermodynamics through longwave  
246 absorption and emission, and the shortwave absorption (Fovell et al. 2016).

247 For this experiment the WSM6 and WDM6 microphysics schemes will be compared. In this  
248 study, I will examine the sensitivities of tropical cyclone boundaries and rotating convection to  
249 single- and double-moment microphysics schemes. Some questions to answer include: How mi-  
250 crophysics parameterization affects the distribution, structure, and longevity of rotating and non-  
251 rotating convection in tropical cyclones and, do single and double moment microphysics schemes  
252 alter how tropical cyclone boundaries form and change over time?

### 253 *c. Planetary boundary layer sensitivity in numerical weather prediction*

254 The planetary boundary layer (PBL) is customarily divided into two layers, the surface layer  
255 (constant-flux layer) and the mixed layer (Kepert 2012). In reality, there is no distinct division  
256 between these layers, though the surface layer typically occupies the lowest tenth of the boundary  
257 layer. In the WRF model, the surface layer is governed by the surface layer scheme and the mixed  
258 layer is governed by the PBL scheme. Because there is no distinct division between these two  
259 layers, the PBL scheme must satisfy physics both in the surface and mixed layers depending on the  
260 depth of the surface layer in the model. PBL schemes parameterize vertical mixing and diffusion  
261 due to eddy mixing in numerical weather models. In most cases, the grid spacing in weather

262 models is not fine enough to resolve turbulent mixing and, therefore, this must be parameterized.  
263 There are three major types of PBL schemes: non-local, local, and hybrid.

264 Non-local schemes use first-order closure allowing for mixing between all the layers in the  
265 boundary layer. In this thesis, the Yonsei University (YSU) PBL (Hong et al. 2006) parame-  
266 terization will be used to represent non-local PBL schemes. The YSU scheme uses K-profile  
267 parameterization (KPP). In KPP schemes, the PBL depth plays a crucial role as it can directly  
268 influence mixing depth, and the magnitude, and, height of maximum heating (Kepert 2012). The  
269 advantages of the YSU scheme are that it can accurately simulate deep vertical mixing in buoyancy  
270 driven PBLs and shallower mixing in strong wind environments (Hong and Lim 2006). The PBL  
271 heights are determined by where the bulk Richardson number exceeds zero. The YSU scheme  
272 tends to overdeepen the PBL in deep convective environments, which often results in too much  
273 dry air near the surface (Coniglio et al. 2013). As tropical cyclones tend to be moist, this drawback  
274 is unlikely.

275 Local schemes use a higher-order closure than non-local schemes allowing for mixing to only  
276 occur between adjacent layers. In this thesis, an improved version of the Mellor–Yamada turbu-  
277 lence closure (MYNN3) model (Nakanishi and Niino 2009) will be used to represent local PBL  
278 schemes. Turbulent kinetic energy (TKE) parameterization is used in the MYNN3 scheme. The  
279 PBL height is determined by where the TKE falls below a critical value ( $1.0 * 10^{-6} \frac{m^2}{s^2}$ ). MYNN3  
280 uses a second-order closure scheme and can do well at simulating mixed layers and stable bound-  
281 ary layers; however, it has difficulty capturing deep vertical mixing (Nakanishi and Niino 2006).  
282 The advantage of the MYNN3 scheme is that it can depict statically stable boundary layers well,  
283 which is not particularly advantageous in the environment of a tropical cyclone. Yet, the MYNN3  
284 scheme often does not account fully for deep vertical mixing associated with large eddies or coun-  
285 tergradient fluxes, which results in weaker updrafts than observed (Nakanishi and Niino 2006).

286 Hybrid schemes use a combination of local and non-local mixing to parameterize turbulent  
287 motions in the PBL. In this study, the asymmetric convective model version 2 (ACM2) will be  
288 used to represent hybrid PBL schemes (Pleim 2007a). ACM2 combines the original non-local  
289 ACM with an eddy diffusion such that this scheme uses first-order closure for upward fluxes  
290 (much like a non-local PBL schemes would) and downward fluxes extend from each layer to each  
291 immediately underlying layer (much like local PBL schemes would). Much like the YSU scheme,  
292 the PBL height is determined by where the bulk Richardson number exceeds 0.25. The advantage  
293 of the ACM2 scheme is that it can depict the vertical profiles of potential temperatures and velocity  
294 in the PBL with greater accuracy than solely local or non-local schemes can (Pleim 2007a). Further  
295 validation of the ACM2 scheme has shown that it is able to support the PBL heights typically seen  
296 in afternoon wind profiler data and radar (Pleim 2007b). Like the YSU scheme, the ACM2 scheme  
297 also tends to overdeepen the PBL in deep convective environments (Coniglio et al. 2013). Very  
298 similar findings to these advantages and disadvantages were seen in Xie et al. (2012), where the  
299 choice of PBL schemes can result in sizable differences in the vertical profiles of temperature,  
300 moisture, and momentum in the boundary layer.

301 Li and Pu (2008) tested the sensitivity of the early rapid intensification of Hurricane Emily  
302 (2005) to microphysics and PBL parameterizations. The local (MYJ) and non-local (YSU)  
303 schemes tested showed a significant difference in intensity between these two schemes, producing  
304 a 19-hPa difference in simulated mean sea level pressure. The main reason for this difference  
305 was that the storms' internal structure, specifically the structure of the eyewall convective heat-  
306 ing distribution, surface latent heat flux, and low-level equivalent potential temperature ( $\theta_e$ ), were  
307 strongly influenced by the PBL schemes. Nolan et al. (2009) evaluated PBL parameterizations in  
308 high-resolution simulations of Hurricane Isabel (2003). The local (MYJ) and the non-local (YSU)  
309 PBL schemes simulated tracks nearly identical to observations and were also able to reproduce

310 a boundary layer with a shallow ( $\sim 600$ -m) well-mixed layer and a much deeper ( $\sim 1000$ -m) ra-  
311 dial inflow layer (Nolan et al. 2009). Finally, in the examination of a tropical cyclone in the Bay  
312 of Bengal, a local (MYJ) PBL scheme produced higher ocean surface fluxes than the non-local  
313 (YSU) PBL scheme (Sateesh et al. 2017). The non-local (YSU) PBL scheme produced a better  
314 simulation with respect to winds and pressure distribution, cloud fraction, and track than the lo-  
315 cal (MYJ) PBL scheme. As stated before, it is likely that the local PBL scheme had difficulty  
316 transporting heat and moisture from the low levels to the upper levels.

317 The YSU, MYNN3, and ACM2 PBL schemes capture the variety of PBL parameterizations  
318 used operationally today in numerical weather prediction models. In this study, I will examine the  
319 sensitivities of tropical cyclone boundaries and rotating convection to these three PBL schemes.  
320 Some questions to answer include: How PBL parameterization affects the distribution, structure,  
321 and longevity of rotating and non-rotating convection in tropical cyclones, and do the PBL schemes  
322 alter how and where tropical cyclone boundaries form and how these boundaries change over time?

#### 323 *d. Tropical cyclone boundaries*

324 Tropical cyclones do not have equal supercell potential everywhere as they tend to cluster near  
325 boundaries. There are two major types of boundaries that have been documented in observations  
326 of landfalling tropical cyclones. The first is areas of convergence of the low-level wind due to fric-  
327 tional differences between the ocean and land (Baker et al. 2009; Green et al. 2011). The second  
328 is baroclinic boundaries due to variations in temperature and moisture (Edwards and Pietrycha  
329 2006). Convergent boundaries tend to enhance shear, while baroclinic boundaries can influence  
330 the distribution of CAPE. The warm and mid-level dry air side of baroclinic boundaries has in-  
331 creased CAPE. Edwards and Pietrycha (2006) suggests four distinct classes of boundaries and the  
332 relation to shear and CAPE that may influence tropical cyclone supercell and tornado potential.

333 The first is the buoyancy-limiting case, such that there is supportive vertical shear profiles on both  
334 sides, but sufficient CAPE only on one side of a boundary. The second is the shear-limiting case,  
335 such that there is supportive CAPE on both sides, but favorable shear on one side of the boundary.  
336 The third is the overlapping case, where there is supportive CAPE on one side and supportive  
337 vertical shear on the other side of a boundary. The last class is the null group, which would have  
338 no apparent organization of shear and CAPE. These four distinct classes of boundaries are likely  
339 to promote different risks as they relate to location of convection in tropical cyclones.

340 There is a stark difference in friction over the ocean and over land. This friction can have a large  
341 impact on the low-level winds in tropical cyclones. Powell and Houston (1998) suggested that  
342 changes in surface wind speed from ocean to land may be as large as  $8-10 \frac{m}{s}$  across horizontal  
343 distances of about 10 km. The winds around a tropical cyclone can be approximated as in gradient  
344 wind balance, which is a balance between the pressure gradient force (PGF), the centrifugal force  
345 ( $\frac{mv^2}{r}$ ), and the Coriolis force ( $2\Omega * v * \sin(latitude)$ ). From gradient wind balance, drastic deceler-  
346 ation of the wind also has impacts on the wind direction. As the wind decelerates due to friction,  
347 it is deflected towards the center of the tropical cyclone as the centrifugal force is a function of the  
348 square of the velocity ( $v^2$ ); thus, it becomes smaller faster than the Coriolis force which is only a  
349 function of velocity ( $v$ ) while the PGF remains the same. In observations of Hurricane Ivan (2004),  
350 Baker et al. (2009) showed near surface wind speed changes from ocean to land of  $2-4 \frac{m}{s}$ , which is  
351 not as large as what was proposed in Powell and Houston (1998). Baker et al. (2009) reported that  
352 it seemed plausible that rapidly moving supercells could experience drastically different low-level  
353 wind profiles within spans of a few kilometers in tropical cyclones during landfall. The change in  
354 wind speed and direction due to friction results in increased low-level shear (increased helicity),  
355 which climatological studies suggest is often associated with more frequent tornadoes (Markowski  
356 et al. 2003) and stronger mesocyclones (Baker et al. 2009)

357 The formation of baroclinic boundaries can happen through a variety of processes in the tropical  
358 cyclone envelope. Vescio et al. (1996) first noted that midlevel dry air intrusions have the po-  
359 tential to substantially alter the thermodynamic structure, which can influence tornado outbreaks  
360 and generate baroclinic boundaries in the tropical cyclone environment. Dry air intrusions into  
361 the tropical cyclone can result in local warming and, therefore, baroclinic boundaries (Edwards  
362 and Pietrycha 2006). Curtis (2004) found that tropical cyclones associated with tornado outbreaks  
363 exhibited three noteworthy environmental details. Tropical cyclones with tornado outbreaks had:  
364 lower LCLs, more moisture from the surface to 900 hPa, and, more dry air above 700 hPa, which  
365 is indicative of dry air intrusions, than the tropical cyclones that did not produce tornado outbreaks  
366 or the null cases. The lower LCL height is consistent with the both buoyancy-limiting case from  
367 Edwards and Pietrycha (2006) and the findings from Rasmussen and Blanchard (1998) who noted  
368 that the LCL height for soundings associated with tornadoes were significantly lower than for  
369 soundings associated with only supercells or even non-supercells across the United States. The  
370 resulting temperature and moisture differences caused by midlevel dry air intrusion creates baro-  
371 clinic boundaries which can act as a catalyst for tornado outbreaks in tropical cyclones (Curtis  
372 2004).

373 Baroclinic boundaries have been documented in both observations (Edwards and Pietrycha  
374 2006) and in model simulations (Green et al. 2011; Card 2019) of tropical cyclones. Dry slots  
375 can lead to the formation of baroclinic boundaries due to differential surface heating within the  
376 tropical cyclone rainband region (Edwards and Pietrycha 2006). Relatively cloud-free areas be-  
377 tween rainbands can support a few degrees Celsius of diabatic surface heating (Card 2019). The  
378 asymmetric surface warming can act to locally magnify CAPE and contribute to supercell mainte-  
379 nance (Edwards 2012).

380 Boundaries like those due to frictional differences between land and ocean surfaces and baro-  
381 clinic gradients caused by gradients in temperature and/or moisture can help convection develop  
382 and mature near the coast during tropical cyclone landfall. Dry air intrusions can also act to in-  
383 crease convective instability invigorating convection and helping develop rotating convection in  
384 localized areas.

### 385 **3. Questions and hypotheses**

386 The purpose of this proposal is to investigate the role of microphysics and planetary boundary  
387 layer parameterizations on the structure, distribution, and development of rotating and non-rotating  
388 convection in tropical cyclones, as well as TC-scale boundaries

389 The first part of this study will investigate WSM6 and WDM6 microphysics parameterizations.  
390 The goal will be to determine the effects of microphysics parameterization on the distribution,  
391 structure, and longevity of convection in tropical cyclones. The second part will investigate local,  
392 non-local, and hybrid PBL parameterizations. As with microphysics, the goal is to determine  
393 the effects of PBL parameterizations on the distribution, structure, and longevity of convection in  
394 tropical cyclones. The third part will investigate boundaries at the tropical cyclone, or sub-tropical  
395 cyclone, scale and the impact on convection. The goal will be to understand how these boundaries  
396 form and affect the distribution of convection, and investigate how these boundaries change over  
397 time. The last part will focus on comparing local frictional effects that drive the weakening of the  
398 wind at the surface and compare those to the winds above the boundary layer to investigate if this  
399 can act to generate additional low-level helicity during tropical cyclone landfall.



400 **Question 1: How does varying the microphysics and planetary boundary layer parameteri-**  
401 **zations affect tropical cyclone convection? Does this choice effect the distribution, structure,**  
402 **and longevity of rotating and non-rotating convection?**

403 *Hypothesis for question 1:* Double and single moment microphysics parameterizations in warm  
404 rain processes have an effect on the development, structure, and longevity of tropical cyclone  
405 convection. The limited number of CCN in WDM6 is likely to limit spurious light precipitation.  
406 The limited CCN is also likely to reduce the amount of ice and graupel in the tropical cyclone.  
407 Reduction of graupel in the cloud would result in a more intense storm (McFarquhar et al. 2006;  
408 Fovell et al. 2009). The largest impact from the microphysics schemes will come from the diabatic  
409 heating and cooling, and the resulting consequences to the convective organization of the tropical  
410 cyclones.

411 Planetary boundary layer parameterizations act to vertically mix heat, moisture, and momentum,  
412 which affects the development and structure of tropical cyclone convection. The YSU scheme is  
413 likely to result in too much dry air at the surface near deep convective cells. It is expected that  
414 the MYNN3 PBL scheme will not fully account for the deep vertical mixing associated with large  
415 eddies in the tropical cyclone boundary layer, thus underestimating the vertical transport of heat,  
416 moisture, and momentum. This underestimate will likely result in less intense convection and re-  
417 sult in less prominent tropical cyclone scale boundaries between the rainbands. ACM2 is non-local  
418 for upward fluxes and is likely to experience the same drawbacks as the YSU scheme, resulting in  
419 over deepening of the PBL. The ACM2 PBL scheme is likely to simulate realistic vertical temper-  
420 ature and wind profiles due to the combination of local and non-local fluxes. For these reasons,  
421 the choice of microphysics and PBL parameterization is likely to affect the distribution, structure,  
422 and longevity of rotating and non-rotating convection.

423 **Question 2: Are there boundaries at the tropical cyclone, or sub-tropical cyclone, scale that**  
424 **help develop or intensify convection locally? Is there a link between dry air intrusion and**  
425 **the development baroclinic boundaries, and does this affect rotating convection? How do**  
426 **baroclinically- and convergence-forced boundaries change over time? How do the choice of**  
427 **microphysics and PBL parameterization affect these boundaries?**

428 *Hypothesis for question 2:* Boundaries, like those due to frictional differences between land and  
429 ocean surfaces and baroclinic gradients, helps convection develop and mature near the coast dur-  
430 ing tropical cyclone landfall. Dry air intrusions act to increase convective instability invigorating  
431 convection locally near these boundaries. Dry air intrusion can also result in clearing between rain-  
432 bands leading to the development of surface baroclinic boundaries due to enhanced insolation, as  
433 seen in Card (2019). Both convergent and baroclinic boundaries help develop and enhance upward  
434 vertical motion in localized areas, causing clustering of convection. Microphysics parameteriza-  
435 tions, particularly the lack of spurious reflectivity (i.e., clearing) will result in stronger and more  
436 frequent baroclinic boundaries with the WDM6 microphysics scheme. PBL parameterizations can  
437 also affect boundaries in tropical cyclones since they govern the vertical transport of momentum,  
438 moisture, and temperature. The local PBL scheme (MYNN3) will have difficulty transporting heat  
439 and moisture fluxes from the surface to the layers above. The low-level distribution of heat and  
440 moisture will affect the formation, and intensity, of baroclinic boundaries and CAPE. Momentum  
441 differences between the PBL schemes will result in less vertical shear in the schemes that do not  
442 propagate the effects of surface friction aloft.

443 **Question 3: Does the rate of weakening at the surface (due to frictional effects), compared**  
444 **to above the boundary layer, generate additional low-level helicity in localized areas during**  
445 **landfall?**

446 *Hypothesis for question 3:* Friction at the surface during landfall will act to weaken near surface  
447 winds faster than winds aloft, generating additional low-level helicity in localized regions over  
448 land. This increased low-level helicity (0–3-km) can greatly increase the likelihood of tornado-  
449 genesis at landfall from an increase in mesocyclone strength. Prior studies have suggested that  
450 development of tropical cyclone supercells is linked to the increase in friction as cells transition  
451 from the ocean to land, acting to increase low-level helicity (Gentry 1983; Baker et al. 2009).  
452 Edwards (2012) noted that as tropical cyclones move inland, the wind profiles do not weaken uni-  
453 formly, which can generate additional low-level helicity. This phenomenon was shown in both  
454 hurricanes Beryl (1994) and Ivan (2004). Weakening at the surface due to friction will locally  
455 increase vertical wind shear over land, generating stronger mesocyclones as rotating convection  
456 makes landfall in the tropical cyclone.

## 457 **4. Methodology**

### 458 *a. Model setup*

459 For this research, the Advanced Research WRF version 4.1 will be used in both the static and  
460 vortex following nest configurations. To efficiently use computing resources, an adaptive time step  
461 will also be utilized in all simulations. Each storm will be simulated in two separate steps, a 9-km  
462 run and then a separate 3-km run with a 1-km vortex following nest.

463 First, a 9-km horizontal grid spacing simulation (Domain 1) will be used to provide the initial  
464 and boundary conditions to the higher-resolution, and vortex-following, nests in the second set of

465 simulations [Figs. 6a (350 X 300 gridpoints) and 7a (300 X 350 gridpoints)]. Domain 1 is run  
466 from 0000 UTC 24 August through 1200 UTC 27 August for Hurricane Harvey (2017) and 1200  
467 UTC 8 September through 0000 UTC 12 September for Hurricane Irma (2017). These times allow  
468 24h for the model to spin up prior to using it as initial and boundary conditions for the second set  
469 of simulations. The ERA5 (Copernicus Climate Change Service (C3S) 2019) is used for the initial  
470 and boundary conditions for the 9-km domain at three-hourly intervals. To help the simulation  
471 develop the storms' intensity and convection faster, the cumulus parameterization scheme New  
472 Tiedtke (Zhang and Wang 2017) was used. For consistency, the 9-km domain is run in multiple  
473 configurations covering all the combinations of microphysics and PBL parameterizations to be  
474 tested in the second set of simulations. All of the 9-km simulations have a 10-hPa model top with  
475 50 vertical levels.

476 The 3-km static domain (Domain 2) in the second set of simulations will use the 9-km simulation  
477 as initial and boundary conditions [Figs. 6b (750 X 600 gridpoints) and 7b (600 X 750 gridpoints)].  
478 The 1-km vortex following domain (Domain 3) is nested within Domain 2 [Figs. 6b and 7b (901 X  
479 901 gridpoints)]. Convective processes will be explicitly resolved in the 3-km and 1-km domains,  
480 therefore no convective parameterization will be used. Both Domain 2 and 3 will have a 50-hPa  
481 model top with 50 vertical levels. In each of these simulations, the microphysics and PBL schemes  
482 will be the only parameterizations varied. Domains 2 and 3 will be run from 0000 UTC 25 August  
483 through 1200 UTC 27 August for Hurricane Harvey (2017) and 1200 UTC 9 September through  
484 0000 UTC 12 September for Hurricane Irma (2017). This timing will provide a 12-h adjustment  
485 period from model start to the analysis times.

486 The WRF model configuration will be set following previous tropical cyclone studies (e.g.,  
487 Gentry and Lackmann 2010; Sun and Barros 2012, 2014; Lackmann 2015; Carroll-Smith 2018).  
488 All three domains for each simulation will use: the updated Rapid Radiative Transfer Model

489 (RRTMG; Iacono et al. 2008) longwave and shortwave radiation schemes; the revised National  
490 Center for Atmospheric Research (NCAR) fifth-generation Mesoscale Model (MM5) Monin-  
491 Obukhov (Jiménez et al. 2012) surface layer parameterization; and, the Noah land surface model  
492 (Chen and Dudhia 2001). To improve the tropical cyclone surface fluxes the “isftcflx” option is  
493 activated such that Donelan and Garratt formulations are used to calculate the surface moist en-  
494 thalpy and momentum exchange coefficients in the surface layer (Lackmann 2015). Individual  
495 model parameterizations for each domain are located in Table 1.

#### 496 *b. Model track and intensity verification*

497 Figure 8a shows that the track of Hurricane Harvey (2017) in the 9-km simulation was very  
498 similar to the observed track in the Atlantic Best Track (Landsea and Franklin 2013) from 0000  
499 UTC 24 August through 1200 UTC 27 August. Early in the 9-km simulation, the track of Harvey  
500 is slightly too far south between 0600 UTC 24 August through 0000 UTC 25 August. As a result,  
501 the 1-km simulations, shown in Figure 8b, are initialized slightly too far south and continue on  
502 a track south of that observed in the Best Track. The 1-km simulations penetrated further inland  
503 before beginning to turn back out to sea.

504 Figure 9a shows that the 9-km WRF simulation of Hurricane Harvey was initialized at a similar  
505 intensity to the observed Harvey, but remained much weaker than observed in the Atlantic Best  
506 Track from 0300 UTC 24 August through 1200 UTC 26 August. There is a divergence of the  
507 model simulations at 0000 UTC 25 August where the model simulations using the MYNN3 PBL  
508 scheme do not strengthen as rapidly as the other simulations. The most intense run is the model  
509 simulation using WDM6 microphysics and YSU PBL parameterizations, reaching a minimum  
510 sea level pressure of 965 hPa compared to the observed minimum sea level pressure of 942 hPa.

511 Figure 9b shows that the 1-km WRF simulations were initialized at a weaker intensity than what

512 was observed in the Best Track. The 1-km simulation using the MYNN3 PBL scheme also showed  
513 a weaker intensity compared to the other simulations over the first 24 h of the simulation.

514 Figure 10a shows that the track of Hurricane Irma (2017) in the 9-km simulation was very similar  
515 to the observed track in the Atlantic Best Track from 1200 UTC 8 September through 0000 UTC  
516 12 September. Like the 9-km simulations, the 1-km simulations, shown in Figure 10b, produce  
517 very similar tracks to the Best Track along the entire analysis time.

518 Figure 11a shows that the 9-km WRF simulation of Hurricane Irma was initialized and remained  
519 much weaker than the observed intensity from the Atlantic Best Track. This discrepancy in the  
520 intensity is not surprising since coarse resolution models such as the ERA5 ( $0.25^\circ \times 0.25^\circ$ ), which  
521 was used to initialize the 9-km WRF, had a minimum sea level pressure at 1200 UTC 8 September  
522 of about 960 hPa. Though the simulations of Hurricane Harvey did better with respects to intensity  
523 than the Hurricane Irma simulation, this is most likely due to the fact that Hurricane Harvey was  
524 much weaker in observations at the start of the simulation compared to Hurricane Irma. The 9-  
525 km WRF simulations have Hurricane Irma maintaining intensity between 955 and 970 hPa before  
526 beginning to weaken after 0000 UTC 11 September. It is seen again that the two weakest runs  
527 were the simulations using the MYNN3 PBL scheme. Figure 11b shows that the 1-km simulations  
528 of Hurricane Irma were stronger than the 9-km simulations; however, the 1-km simulation did  
529 not become as intense as the observations from the Best Track. Like the 9-km simulation, the  
530 WDM6–MYNN3 1-km simulation produced a less intense storm.

531 Although the simulations of both hurricanes Harvey and Irma (2017) are weaker than the ob-  
532 served storms and took slightly different tracks, the simulations show similar tracks and intensities  
533 to each other. These similarities will allow for a clean comparison between the differences pro-  
534 duced by changing the microphysics and boundary layer parameterizations.

535 *c. Diagnostics*

536 A series of analyses will be used to investigate the interactions between microphysics and plan-  
537 etary boundary layer parameterizations on the development and structure of tropical cyclone con-  
538 vection. The first analysis will focus on the distribution and structure of rotating and non-rotating  
539 convection. The second analysis will focus on the location and evolution of frictional and baro-  
540 clinic boundaries in the tropical cyclones, and how these boundaries vary with different micro-  
541 physics and planetary boundary layer schemes. A mesoscale analysis will be used to investigate  
542 the interactions between various types of boundaries and convection in the rainbands of tropical  
543 cyclones. This study is unique in that it plans to investigate the interactions between microphysics  
544 and planetary boundary layer parameterizations on the development, evolution, and structure of  
545 both tropical cyclone convection and boundaries during landfall.

546 1) DISTRIBUTION AND STRUCTURE OF ROTATING AND NON-ROTATING CONVECTION

547 To investigate the interaction between microphysics and PBL parameterizations on the distribu-  
548 tion of rotating and non-rotating convection, an analysis similar to the techniques of Card (2019)  
549 and Carroll-Smith et al. (2019) will be done on the 1-km WRF domain. First, individual convec-  
550 tive cells will be identified using local maxima in model reflectivity exceeding the 99.9<sup>th</sup> percentile  
551 across all the hours of a simulation. The identified cells will be referred to as rotating convective  
552 cells if the 0–3-km updraft helicity exceeds the 99.95<sup>th</sup> similar to Carroll-Smith et al. (2019). The  
553 identified cells that had values of 0–3-km updraft helicity less than or equal to zero and that did  
554 exceed the 99.9<sup>th</sup> percentile in updraft velocity, will be referred to as non-rotating convective cells.  
555 Based on these criteria, the non-rotating convective cells have no updraft helicity but strong updraft  
556 velocities, while rotating cells have large updraft helicity. These percentile values for hurricanes  
557 Harvey and Irma (2017) can be seen in Tables 2 and 3, respectively.

558 In these thresholds is where we see the first indication of differences across microphysics and  
559 PBL schemes. The 99.9<sup>th</sup> percentile in model reflectivity, 99.9<sup>th</sup> percentile in updraft velocity, and  
560 the 99.95<sup>th</sup> percentile in 0–3-km updraft helicity were all statistically significantly<sup>1</sup> different across  
561 the micorphysycis and PBL parameterizations. In Tables 2 and 3, the 99.9<sup>th</sup> percentile in model  
562 reflectivity is statistically significantly lower in the simulations using the WDM6 microphysics  
563 scheme compared to WSM6 simulations using the same PBL scheme. This result was expected as  
564 past research has shown that WDM6 tends to produce less spurious reflectivity than WSM6 (Hong  
565 et al. 2010). Tables 2 and 3 also show that the 99.9<sup>th</sup> percentile in updraft velocity and the 99.95<sup>th</sup>  
566 percentile in 0–3-km updraft helicity was statistically significantly lower in the simulations using  
567 the MYNN3 PBL scheme.

568 After identification, the rotating and non-rotating convection distributions across the two micro-  
569 physics and three PBL parameterizations will be examined. The distributions will be examined  
570 with respect to vertical wind shear and with respect to geographic north. Card (2019) and the  
571 storm tracks in Figures 8 and 10 showed that for hurricanes Harvey and Irma (2017), the storm  
572 motion was mostly northerly making the results very comparable to the distributions with respect  
573 to geographic north. In Card (2019), the number of identified rotating storms outnumbered the  
574 identified non-rotating storms by a factor of 2–3 in both Harvey and Irma (2017). In Hurricane  
575 Harvey, most rotating and non-rotating storms occurred directly downshear, with non-rotating  
576 storms generally occurring at more distant radii (Fig. 4). Most of the rotating storms in Hurricane  
577 Irma occurred directly downshear, while most of the non-rotating storms occurred upshear right in  
578 both the NCAR ensemble and in observations (Fig. 5). Once cell types are identified, how varying  
579 the microphysics and planetary boundary layer schemes changes the spatial distribution of rotating  
580 and non-rotating convective cells will be examined.

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<sup>1</sup>Statistical significance is determined via a two sided t-test for the means of two independent samples at the 99% confidence level.



581 Finally, vertical cross sections will be done through a few of the select rotating and non-rotating  
582 cells to examine the vertical structure and spatial distributions of mixing ratios of water vapor,  
583 rain, and ice across the different microphysics parameterizations. Vertical cross sections will also  
584 be done for select rotating and non-rotating convective cells to examine the vertical structure and  
585 boundary layer interactions in the lowest 3 km. Similarly to Didlake and Houze (2009), who did  
586 cross sections of Hurricane Katrina (2005) during the Rainband and Intensity Change Experiment  
587 (RAINEX), these cross sections will be used to investigate how convective scale wind flow and  
588 other convective scale features within the individual cells.

589 The goals of the investigation into the distribution, structure, and longevity of rotating and non-  
590 rotating convection are to investigate Question one.

## 591 2) FRICTIONAL AND BAROCLINIC BOUNDARIES

592 Boundaries, both induced by friction and those induced by baroclinic features will be investi-  
593 gated along the Texas coast in Hurricane Harvey and the Florida coast in Hurricane Irma in the  
594 1-km simulations. Boundaries induced by friction will be identified by diagnosing the wind field  
595 near the coastline, while baroclinic boundaries will be identified using gradients in temperature,  
596 relative humidity, and MUCAPE (most unstable CAPE). Furthermore, vertical cross sections of  
597 these boundaries will help in understanding the depth of these features, how they may change or  
598 move over time, and how they affect convection. In association with Question one, differences in  
599 the boundaries based on the microphysics and PBL parameterizations will be examined.

600 An analysis of the three-dimensional frontogenesis equation (eq. 1) will provide insight into  
601 which terms might be important in tropical cyclone boundary formation:

$$\begin{aligned}
F = & \frac{1}{|\nabla\theta|} \left\{ \frac{\partial\theta}{\partial x} \left[ \frac{1}{C_p} \left( \frac{p_o}{p} \right)^\kappa \left[ \frac{\partial}{\partial x} \left( \frac{dQ}{dt} \right) \right] - \left( \frac{\partial u}{\partial x} \frac{\partial\theta}{\partial x} \right) - \left( \frac{\partial v}{\partial x} \frac{\partial\theta}{\partial y} \right) - \left( \frac{\partial w}{\partial x} \frac{\partial\theta}{\partial z} \right) \right] \right. \\
& + \frac{\partial\theta}{\partial y} \left\{ \frac{1}{C_p} \left( \frac{p_o}{p} \right)^\kappa \left[ \frac{\partial}{\partial y} \left( \frac{dQ}{dt} \right) \right] - \left( \frac{\partial u}{\partial y} \frac{\partial\theta}{\partial x} \right) - \left( \frac{\partial v}{\partial y} \frac{\partial\theta}{\partial y} \right) - \left( \frac{\partial w}{\partial y} \frac{\partial\theta}{\partial z} \right) \right\} \\
& \left. + \frac{\partial\theta}{\partial z} \left\{ \frac{p_o^\kappa}{C_p} \left[ \frac{\partial}{\partial z} \left( p^{-\kappa} \frac{dQ}{dt} \right) \right] - \left( \frac{\partial u}{\partial z} \frac{\partial\theta}{\partial x} \right) - \left( \frac{\partial v}{\partial z} \frac{\partial\theta}{\partial y} \right) - \left( \frac{\partial w}{\partial z} \frac{\partial\theta}{\partial z} \right) \right\} \right\} \quad (1)
\end{aligned}$$

602 where theta ( $\theta$ ) is the potential temperature,  $C_p$  is the specific heat at constant pressure  
603 ( $1006 \frac{J}{kgK}$ ),  $p$  is the pressure,  $p_o$  is a reference pressure (1000hPa),  $\kappa$  is a constant [ $\frac{R}{C_p}$ , 0.286],  
604  $Q$  is diabatic heating,  $u$  is the zonal wind, and  $v$  is the meridional wind. The three-dimensional  
605 frontogenesis equation can be broken up into four major components; 1) the diabatic terms (eq.  
606 2), 2) the deformation terms (eq. 3), 3) the tilting terms (eq. 4), and 4) the vertical divergence term  
607 (eq. 5). The tilting term (eq. 4) can not generate gradients in potential temperature, it can only  
608 transform vertical gradients into the horizontal and, therefore, will not be included in the analysis  
609 of the three-dimensional frontogenesis equation.

$$Diabatic = \frac{1}{C_p} \left( \frac{p_o}{p} \right)^\kappa \left[ \frac{\partial}{\partial x} \left( \frac{dQ}{dt} \right) \right] + \frac{1}{C_p} \left( \frac{p_o}{p} \right)^\kappa \left[ \frac{\partial}{\partial y} \left( \frac{dQ}{dt} \right) \right] + \frac{p_o^\kappa}{C_p} \left[ \frac{\partial}{\partial z} \left( p^{-\kappa} \frac{dQ}{dt} \right) \right] \quad (2)$$

$$\begin{aligned}
Deformation = & - \left( \frac{\partial u}{\partial x} \frac{\partial\theta}{\partial x} \right) - \left( \frac{\partial v}{\partial x} \frac{\partial\theta}{\partial y} \right) \\
& - \left( \frac{\partial u}{\partial y} \frac{\partial\theta}{\partial x} \right) - \left( \frac{\partial v}{\partial y} \frac{\partial\theta}{\partial y} \right) \\
& - \left( \frac{\partial u}{\partial z} \frac{\partial\theta}{\partial x} \right) - \left( \frac{\partial v}{\partial z} \frac{\partial\theta}{\partial y} \right)
\end{aligned} \quad (3)$$

$$Tilting = - \left( \frac{\partial w}{\partial x} \frac{\partial\theta}{\partial z} \right) - \left( \frac{\partial w}{\partial y} \frac{\partial\theta}{\partial z} \right) \quad (4)$$

$$Vertical Divergence = - \left( \frac{\partial w}{\partial z} \frac{\partial\theta}{\partial z} \right) \quad (5)$$

610 To determine the diabatic heating rate ( $\frac{dQ}{dt}$ ) needed for the diabatic term in the three-dimensional  
 611 kinematic frontogenesis equation (eq. 2), I will use the temperature tendency equation (eq. 6) after  
 612 applying the first law of thermodynamics (eq. 7) as in Yanai et al. (1973).

$$\frac{dT}{dt} = \frac{\partial T}{\partial t} + \vec{V} \cdot \nabla T \quad (6)$$

$$\frac{dT}{dt} = -\frac{g}{c_p} w + \frac{1}{c_p} \frac{dQ}{dt} \quad (7)$$

613 Combining equations 6 and 7 and reorganizing produces an equation for the diabatic heating  
 614 rate ( eq. 8).

$$\frac{dQ}{dt} = c_p \left( \frac{\partial T}{\partial t} + \vec{V} \cdot \nabla T + \frac{g}{c_p} w \right) \quad (8)$$

615 This study will look at the diabatic (eq. 2), deformation (eq. 3), and vertical divergence (eq.  
 616 5) terms and preform a scale analysis to determine which terms may play the largest role in the  
 617 tropical cyclone environment. The overarching goals of the investigation into the frictional and  
 618 baroclinic boundaries during landfall are to address Questions two and three.

## 619 **5. Preliminary findings**

### 620 *a. Cell distribution thresholds*

621 Preliminary results from the cell distribution thresholds has provided some interesting results.  
 622 As seen from the 99.95<sup>th</sup> percentile values of the 0–3-km updraft helicity in Tables 2 and 3, the  
 623 MYNN3 PBL scheme produces statistically significantly lower values than either the YSU or  
 624 ACM2 schemes. There are two terms which contribute to 0–3-km updraft helicity, the vertical  
 625 velocity and the 0–3-km helicity. To understand the differences in these two terms we will focus

626 on 1800 UTC 26 August for Hurricane Harvey and 1800 UTC 10 September for Hurricane Irma.  
627 Figures 12 and 13 show the differences in maximum updraft velocity between the MYNN3 PBL  
628 scheme and the other two PBL schemes. The differences are mainly on the convective scale,  
629 showing where there is convection in each simulation there is strong vertical motion. In both  
630 hurricanes Harvey and Irma, large differences in convection are located on the eastern half of the  
631 storms. The location of the convection is not highly spatially correlated between the simulations.

632 Figures 14 and 15 show the differences in 0–3-km helicity between the MYNN3 PBL scheme  
633 and the other two PBL schemes at these times. These differences are larger in spatial scale than  
634 those of the vertical velocities. The differences in the 0–3-km helicity are mainly on the northern  
635 side of the storm in both hurricanes Harvey and Irma, which happens to be the land side of the  
636 storm.

637 The planetary boundary layer height can be used to understand how the simulations are interact-  
638 ing with the land and how deep the vertical mixing is in this area of the storm. This will provide  
639 insight as to why the spatial pattern of 0–3-km helicity may be different over land. Figure 16 shows  
640 the PBL heights for the simulations. For both storms the PBL heights tend to be much deeper over  
641 land in the YSU and ACM2 PBL schemes, compared to the MYNN3 scheme. Future work will  
642 include diagnosing how the depth of the PBL between these simulations affects the low-level  
643 helicity, and continuing to explore the questions posed in Question one.

#### 644 *b. Boundaries*

645 Preliminary results from the frontogenesis equation for Hurricane Harvey at 1800 UTC 26 Au-  
646 gust and Hurricane Irma at 1800 UTC 10 September are shown in Figure 17. For the simulations  
647 using WSM6–YSU parameterizations, both storms show that the diabatic heating term is the lead-  
648 ing term in the full frontogenesis equation. Figure 18 shows the outgoing longwave radiation as a

649 proxy for convection intensity, where colder cloud tops signify more vigorous convection. These  
650 differences in convection are highlighted in the diabatic heating term of the frontogenesis equation  
651 in Figure 17. In the WSM6–YSU and WDM6–YSU simulations, the coldest cloud tops are lo-  
652 cated to the northeast of the center of both hurricanes Harvey and Irma (Fig. 18). In WDM6–YSU  
653 simulation for Hurricane Irma (Fig. 18d), warm cloud tops can be seen surrounding the center of  
654 the storm, unlike in the WSM6–YSU simulation (Fig. 18b).

655 The deformation terms in the frontogenesis equation are strongly tied to horizontal gradients in  
656 potential temperature gradients (Fig. 17). In both hurricanes Harvey and Irma, areas with lower  
657 relative humidity display potential temperatures of about 2 K warmer compared to the moist areas  
658 (Fig. 19), which is very similar to what was seen in Card (2019). There tends to be a larger areal  
659 spread of high relative humidity in the WDM6–YSU simulations if both hurricanes Harvey and  
660 Irma (Fig. 19c and 19d).

661 These preliminary results suggest that in both the simulations of hurricanes Harvey and Irma that  
662 prescribed PBL and microphysics parameterizations will impact frontogenesis. Both the diabatic  
663 heating from differences in convection in the tropical cyclone and the deformation of temperature  
664 gradients in relation to areas of lower relative humidity and wind. Future work will focus on  
665 continuing to explore the questions posed in Questions two and three.

## 666 **6. Timeline of Work**

### 667 **Summer 2020:**

- 668 ● Begin writing introduction, literature review, and methods sections of the dissertation
- 669 ● Continue analysis for all hypothesis

### 670 **Fall 2020:**

671 • Begin writing results chapters

672 • Edit full Ph.D. dissertation

673 **December 2020:**

674 • Defend Ph.D. dissertation.

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TABLE 1. Select WRF model settings and parameterizations.

<b>WRF Model Parameterizations:</b>	<b>9-km Simulation (Domain 1)</b>	<b>3-km Simulation (Domain 2)</b>	<b>1-km Simulation (Domain 3)</b>	
<b>Grid spacing</b>	9 km	3 km	1 km	
<b>Vertical levels (Model top)</b>	50 (10 hPa)	50 (50 hPa)	50 (50 hPa)	
<b>Vortex following (Level)</b>	--	--	Yes (500 hPa)	
<b>Long- and shortwave radiation parameterization</b>	RRTMG	RRTMG	RRTMG	(Iacono et al. 2008)
<b>Land surface model</b>	Noah	Noah	Noah	(Chen and Dudhia 2001)
<b>Surface layer physics</b>	Revised MM5 Monin-Obukhov scheme	Revised MM5 Monin-Obukhov scheme	Revised MM5 Monin-Obukhov scheme	(Jimenez et al. 2012)
<b>Other physics</b>	TC surface fluxes	TC surface fluxes	TC surface fluxes	Donelan Cd + Garratt Ck (Lackmann 2015)
<b>Cumulus parameterization</b>	Tiedtke	--	--	(Zhang et al. 2011)
<b>Microphysics</b>	WSM6, WDM6	WSM6, WDM6	WSM6, WDM6	(Hong and Lim 2006), (Lim and Hong 2010)
<b>Planetary boundary layer</b>	YSU, MYNN3, ACM2	YSU, MYNN3, ACM2	YSU, MYNN3, ACM2	(Hong et al. 2006), (Nakanishi and Niino 2006), (Pleim 2007a,b)



865 TABLE 2. Percentile threshold values for cell type identification for Hurricane Harvey (2017). The model  
 866 reflectivity values for the 99.9<sup>th</sup> percentile, 0–3-km updraft helicity values for the 99.95<sup>th</sup> percentile, and updraft  
 867 velocity values for the 99.9<sup>th</sup> percentile are shown.

<b>WRF Model Runs</b>	<b>Model Reflectivity (dBz)</b>	<b>0–3-km Updraft Helicity (m<sup>2</sup>/s<sup>2</sup>)</b>	<b>Max Updraft Velocity (m/s)</b>
<b>Harvey (2017):</b>			
<b>WSM6—YSU</b>	52.41	32.95	15.13
<b>WSM6—MYNN3</b>	51.24	21.71	11.17
<b>WSM6—ACM2</b>	53.25	37.85	16.80
<b>WDM6—YSU</b>	51.18	33.64	15.33
<b>WDM6—MYNN3</b>	50.61	21.89	11.31
<b>WDM6—ACM2</b>	51.23	33.92	16.42

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 870 velocity values for the 99.9<sup>th</sup> percentile are shown.

<b>WRF Model Runs</b>	<b>Model Reflectivity (dBz)</b>	<b>0–3-km Updraft Helicity (m<sup>2</sup>/s<sup>2</sup>)</b>	<b>Max Updraft Velocity (m/s)</b>
<b>Irma (2017):</b>			
<b>WSM6—YSU</b>	54.74	109.48	17.18
<b>WSM6—MYNN3</b>	53.47	57.62	13.51
<b>WSM6—ACM2</b>	55.31	126.32	19.33
<b>WDM6—YSU</b>	52.34	98.45	20.16
<b>WDM6—MYNN3</b>	51.91	65.43	13.98
<b>WDM6—ACM2</b>	52.59	113.50	18.31

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898 initialized at 1200 UTC 8 September and b) 1-km WRF simulation initialized at 1200 UTC  
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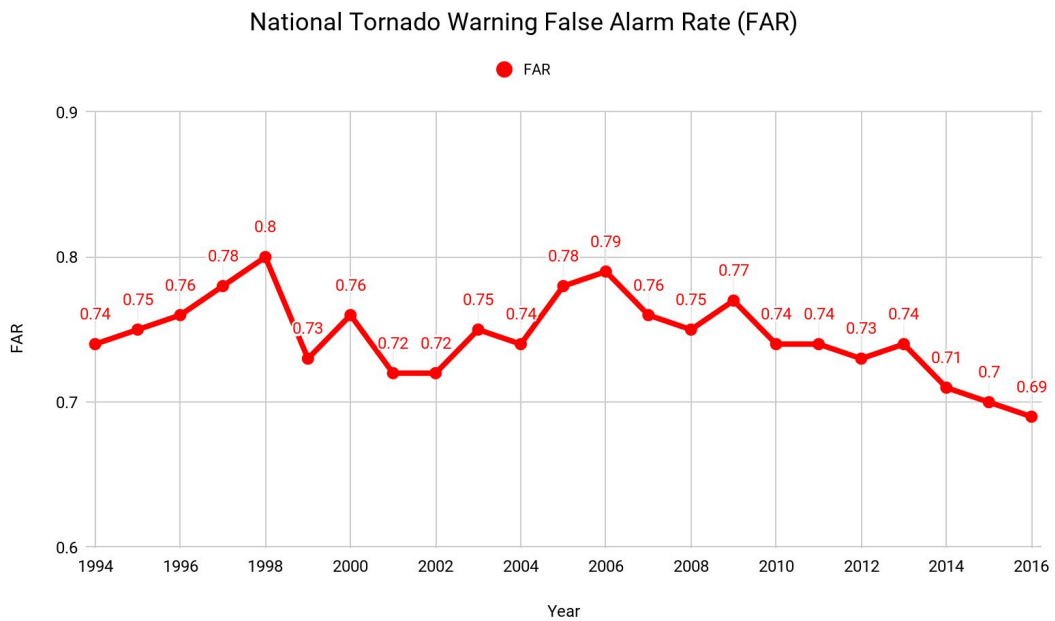
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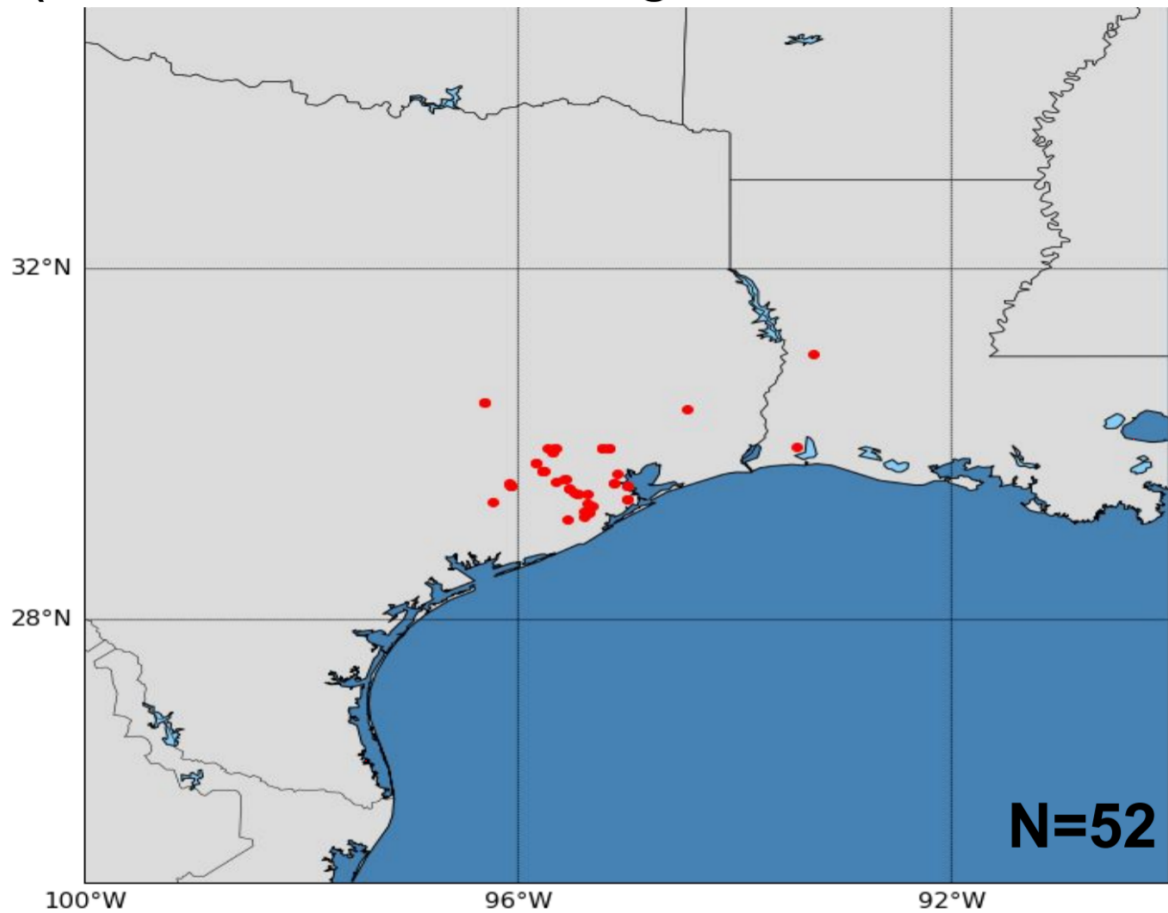
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909	<b>Fig. 15.</b>	0–3-km helicity difference ( $m^2 s^{-2}$ , shaded) for the MYNN3 PBL scheme between the YSU	
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911	<b>Fig. 16.</b>	Planetary boundary layer height ( $m$ , shaded) for the MYNN3, YSU, and ACM2 PBL	
912		schemes in hurricanes Harvey (top) and Irma (bottom) (2017) at 1800 UTC 26 August and	
913		1800 UTC 10 September, respectively. . . . .	62
914	<b>Fig. 17.</b>	WSM6–YSU simulation 1500-m diabatic, deformation, vertical divergence, and the full	
915		frontogenesis equations ( $\frac{K}{skm}$ , shaded) at a) 1800 UTC 26 August for Hurricane Harvey	
916		(2017) and b) at 1800 UTC 10 September for Hurricane Irma (2017). . . . .	63
917	<b>Fig. 18.</b>	Outgoing longwave radiation (K, shaded) for the a) WSM6 and c) WDM6 simulations for	
918		1800 UTC 26 August for Hurricane Harvey (2017) and b) WSM6 and d) WDM6 simulations	
919		at 1800 UTC 10 September for Hurricane Irma (2017). . . . .	64
920	<b>Fig. 19.</b>	1500-m relative humidity (% , shaded), potential temperature (K, dashed), and wind barbs	
921		( $m s^{-1}$ , standard convention) for the a) WSM6 and c) WDM6 simulations for 1800 UTC 26	
922		August for Hurricane Harvey (2017) and b) WSM6 and d) WDM6 simulations at 1800 UTC	
923		10 September for Hurricane Irma (2017). . . . .	65



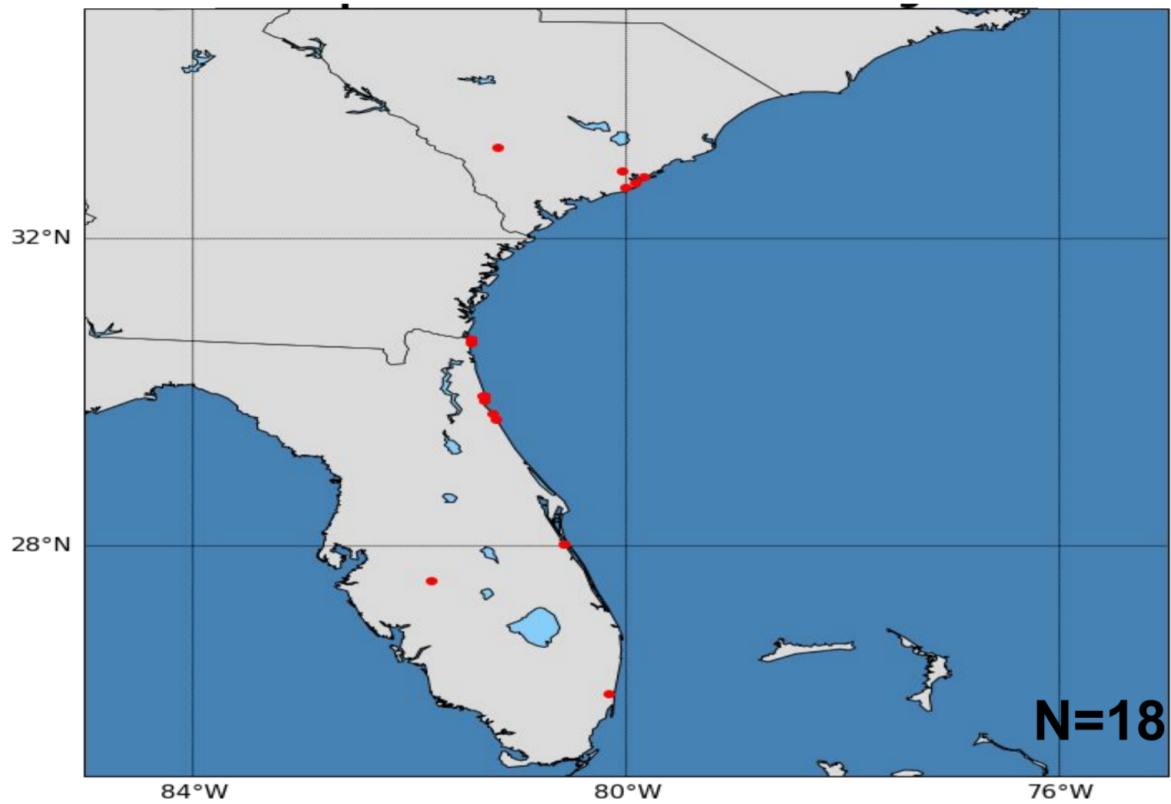
924 FIG. 1. National tornado warning false alarm ratio (FAR) for the United States from the Storm Prediction  
 925 Center (SPC) 1994–2016.

## Harvey Tornado Reports (0000 UTC 2017-08-26 through 1200 UTC 2017-08-27)

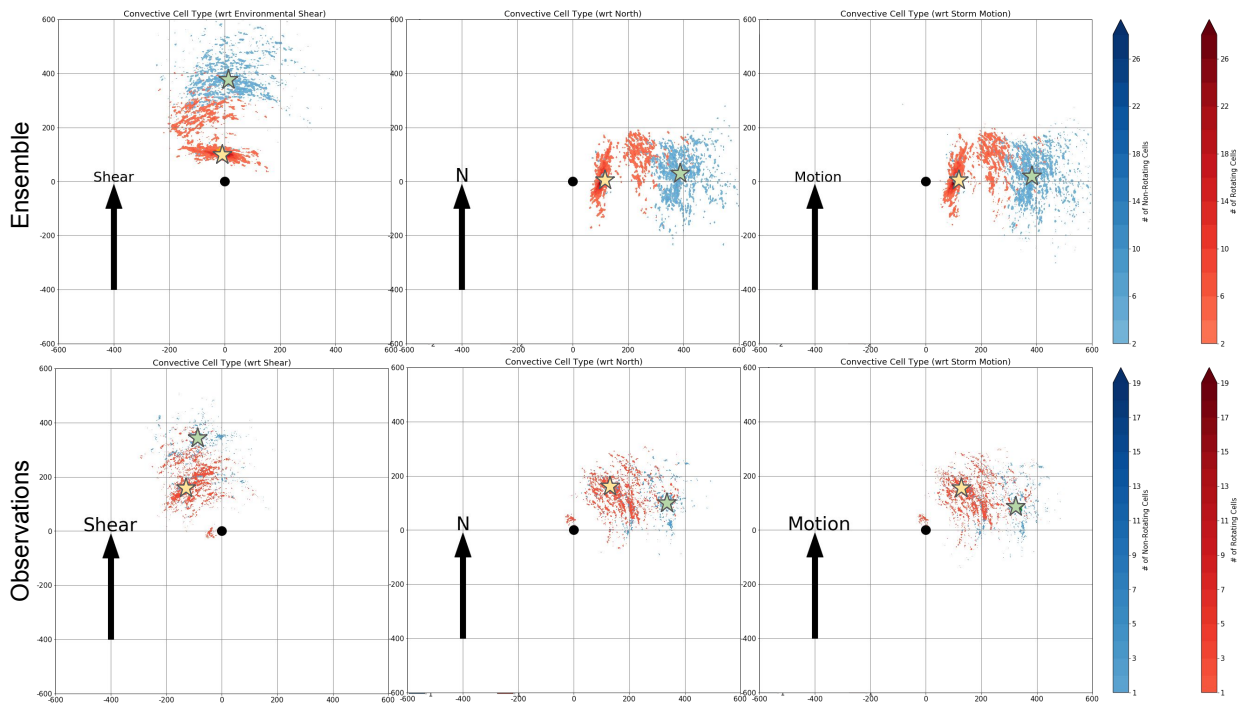


926 FIG. 2. Storm Prediction Center (SPC) tornado reports (red) for 0000 UTC 26 August through 1200 UTC 27  
927 August.

### Irma Tornado Reports (1200 UTC 2017-09-10 through 0000 UTC 2017-09-12)

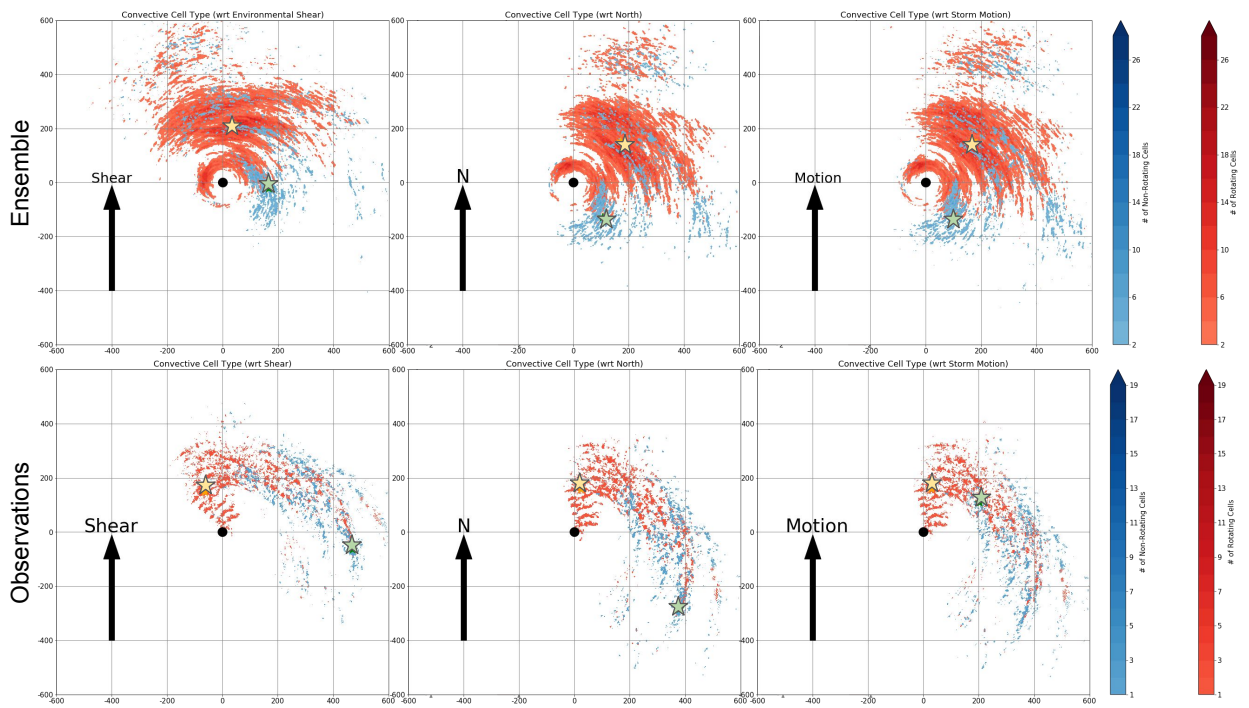


928 FIG. 3. Storm Prediction Center (SPC) tornado reports (red) for 1200 UTC 10 September through 0000 UTC  
929 12 September.

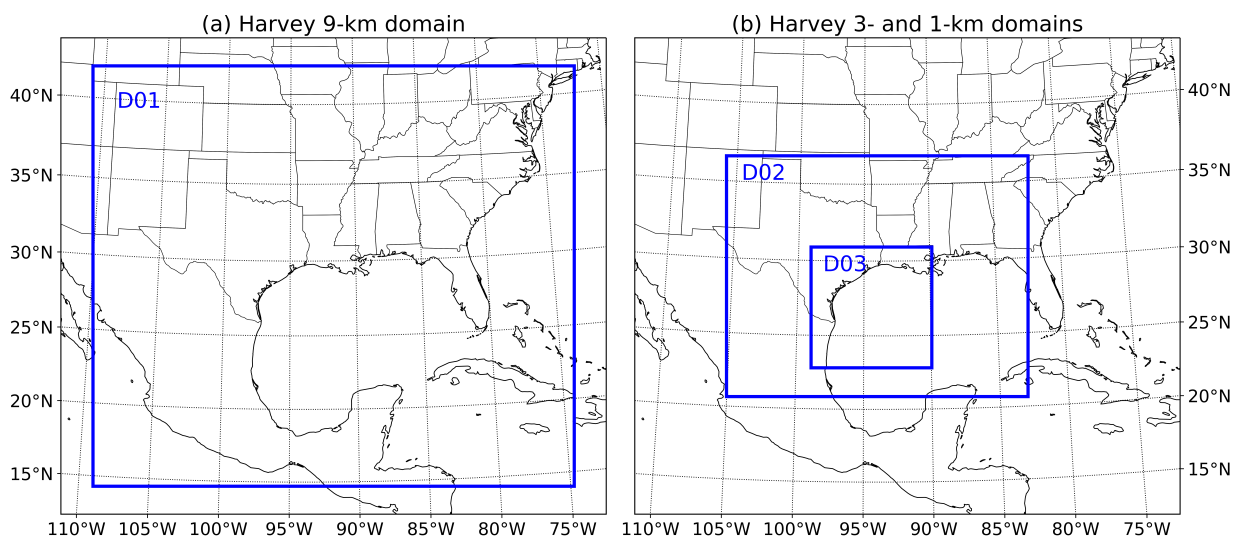


930 FIG. 4. Distribution of rotating (red) and non-rotating (blue) cells in the NCAR ensemble initialized at 0000  
 931 UTC 26 August (top) and observations (bottom) with respect to vertical shear, north, and storm motion from  
 932 0000 UTC 26 August through 1200 UTC 27 August 2017. Center of mass of the rotating cells (star yellow) and  
 933 non-rotating cells (star green). From Card (2019).

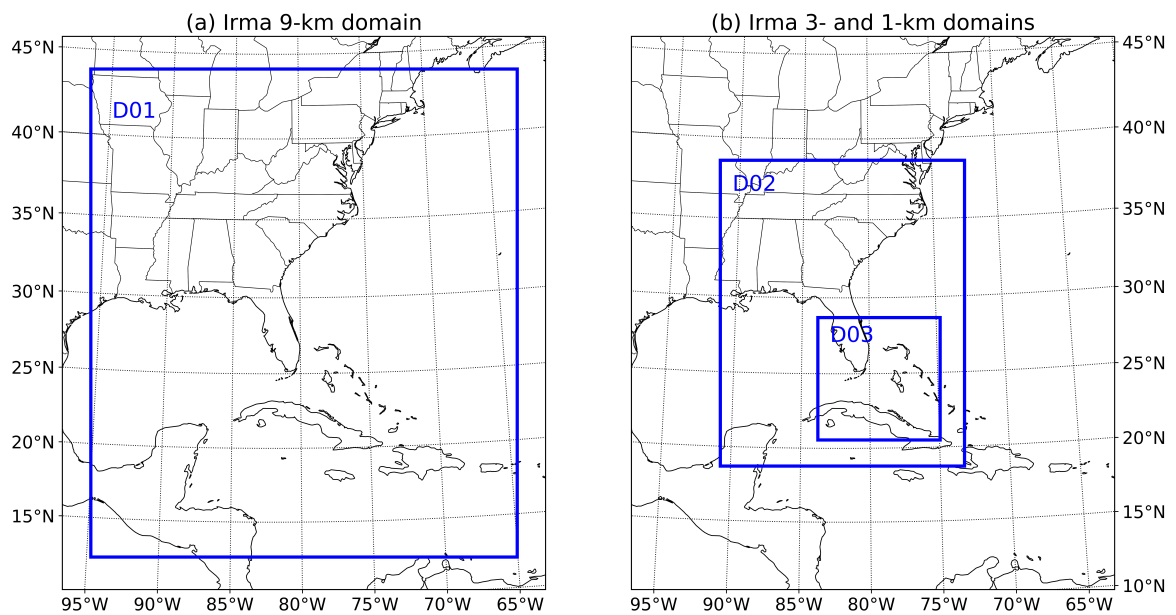




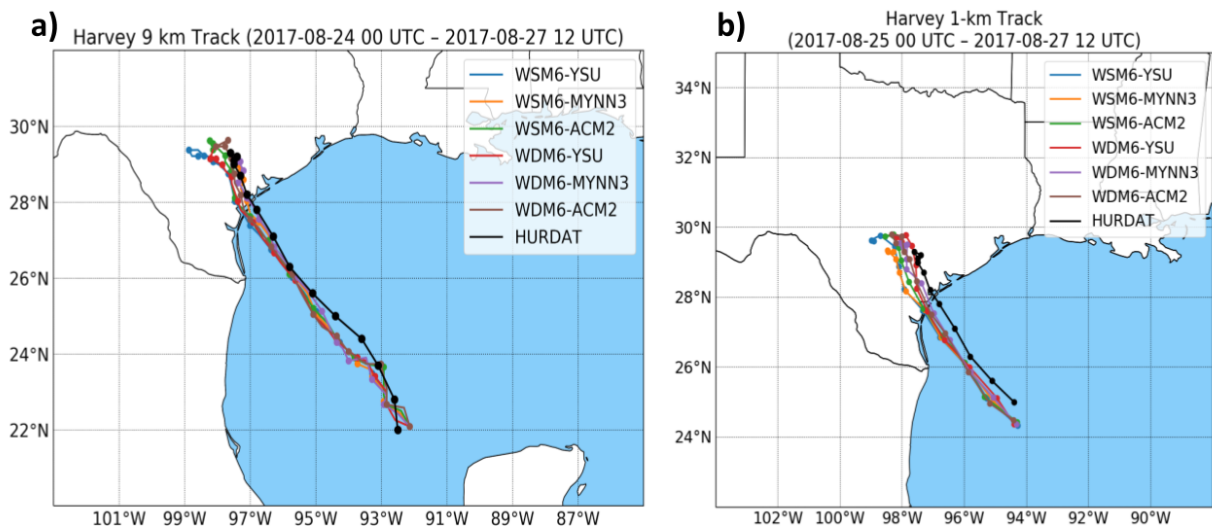
934 FIG. 5. Same as Figure 4, but for Hurricane Irma (2017) from 1200 UTC 10 September through 0000 UTC  
 935 12 September 2017. The NCAR ensemble was initialized at 0000 UTC 10 September. From Card (2019).



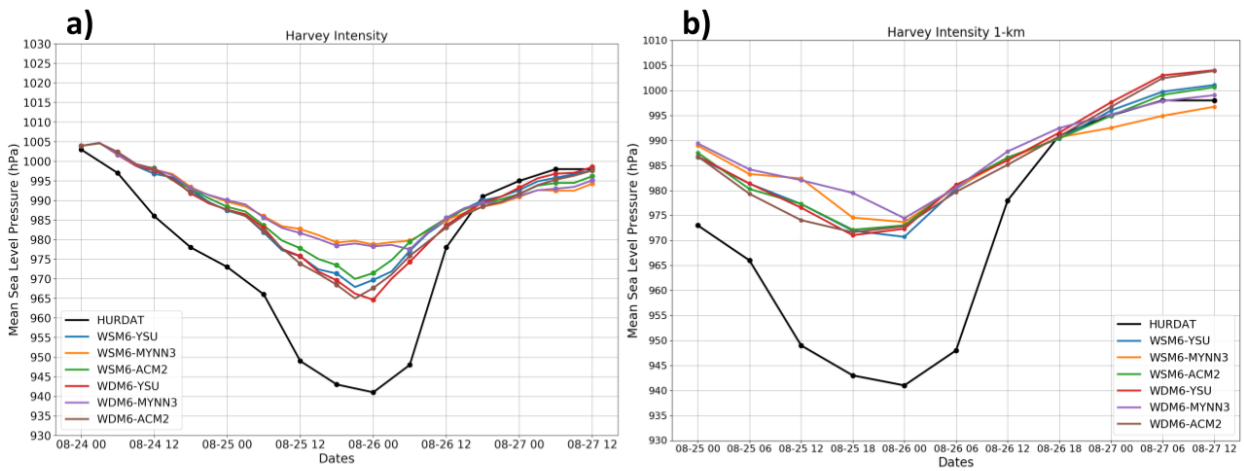
936 FIG. 6. The WRF domains for Hurricane Harvey (2017): a) 9-km domain (D01, 350 X 300 gridpoints)  
 937 and b) 3-km static domain (D02, 750 X 600 gridpoints) with 1-km vortex following domain (D03, 901 X 901  
 938 gridpoints).



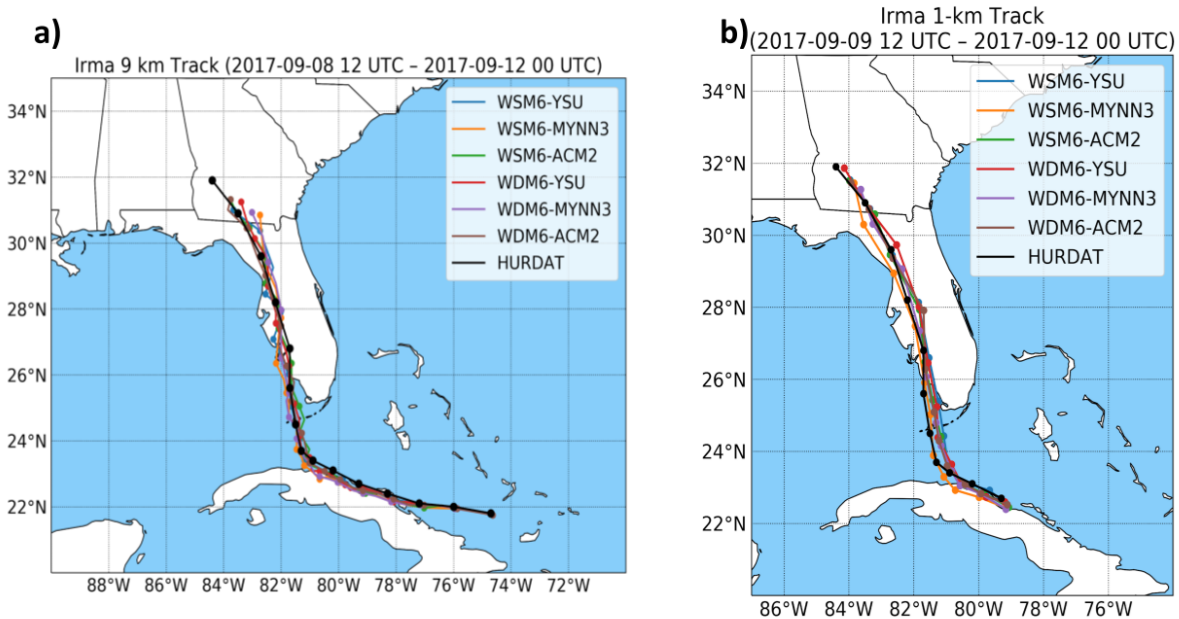
939 FIG. 7. The WRF domains for Hurricane Irma (2017): a) 9-km domain (D01, 300 X 350 gridpoints) and b) 3-  
 940 km static domain (D02, 600 X 750 gridpoints) with 1-km vortex following domain (D03, 901 X 901 gridpoints).



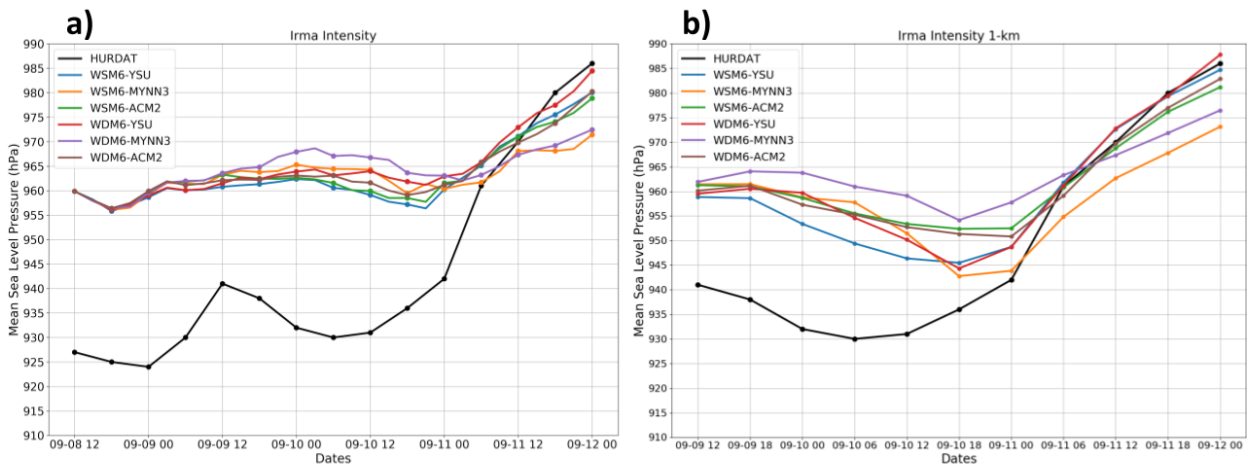
941 FIG. 8. Tropical cyclone tracks using minimum sea level pressure for the a) 9-km WRF simulation initialized  
 942 at 0000 UTC 24 August and b) 1-km WRF simulation initialized at 0000 UTC 25 August, compared to the  
 943 Atlantic Best Track of Hurricane Harvey (2017) every 6 h.



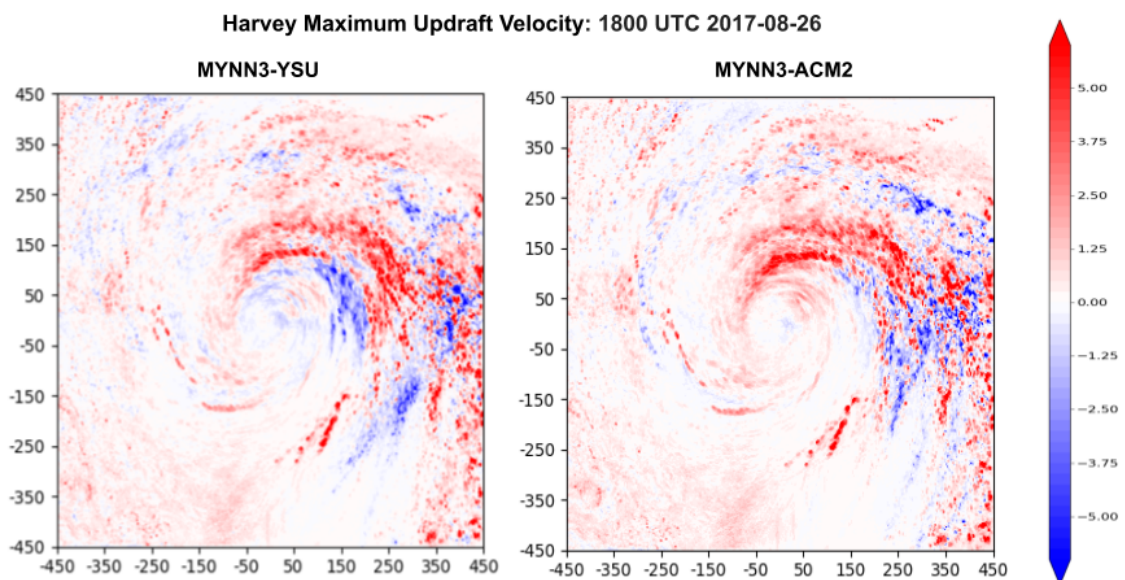
944 FIG. 9. Intensity, in terms of minimum sea level pressure (hPa), for the a) 9-km WRF simulation initialized at  
 945 0000 UTC 24 August and b) 1-km WRF simulation initialized at 0000 UTC 25 August, compared to the Atlantic  
 946 Best Track of Hurricane Harvey (2017) every 6 h.



947 FIG. 10. Tropical cyclone tracks using minimum sea level pressure for the a) 9-km WRF simulation initialized  
 948 at 1200 UTC 8 September and b) 1-km WRF simulation initialized at 1200 UTC 9 September, compared to the  
 949 Atlantic Best Track of Hurricane Irma (2017) every 6 h.

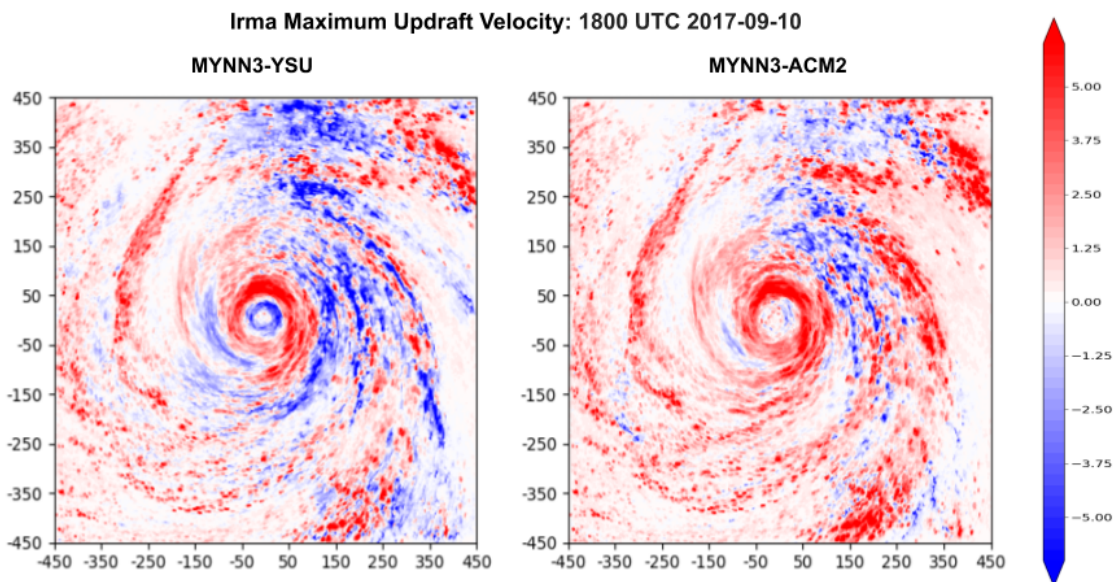


950 FIG. 11. Intensity, in terms of minimum sea level pressure (hPa), from the a) 9-km WRF simulation initialized  
 951 at 1200 UTC 8 September and b) 1-km WRF simulation initialized at 1200 UTC 9 September, compared to the  
 952 Atlantic Best Track of Hurricane Irma (2017) every 6 h.

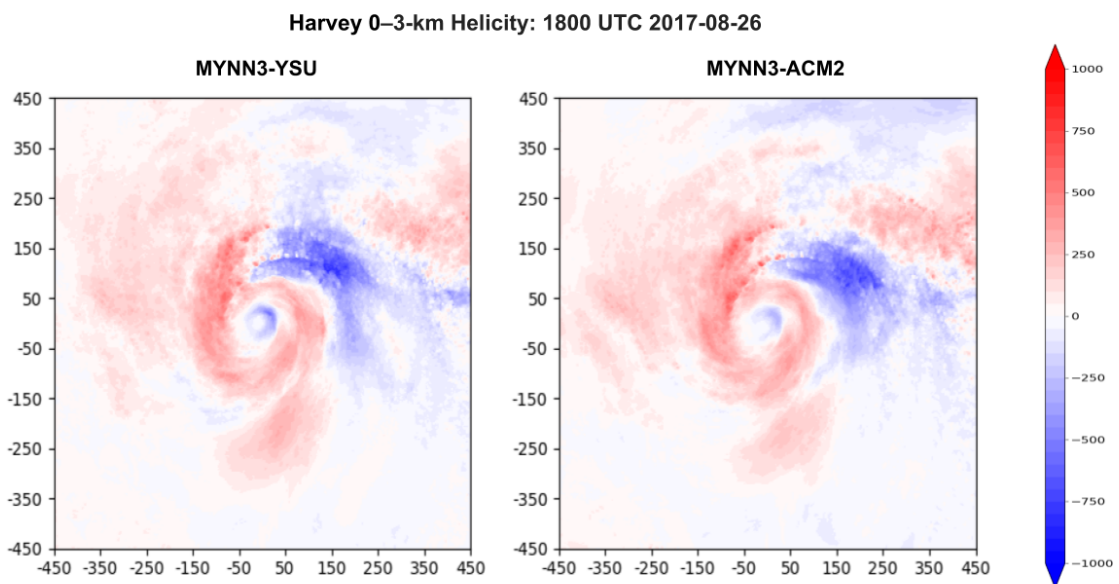


953 FIG. 12. 0–3-km vertical velocity difference ( $m s^{-1}$ , shaded) for the MYNN3 PBL scheme between the YSU  
 954 (left) and ACM2 (right) for Hurricane Harvey (2017) at 1800 UTC 26 August.

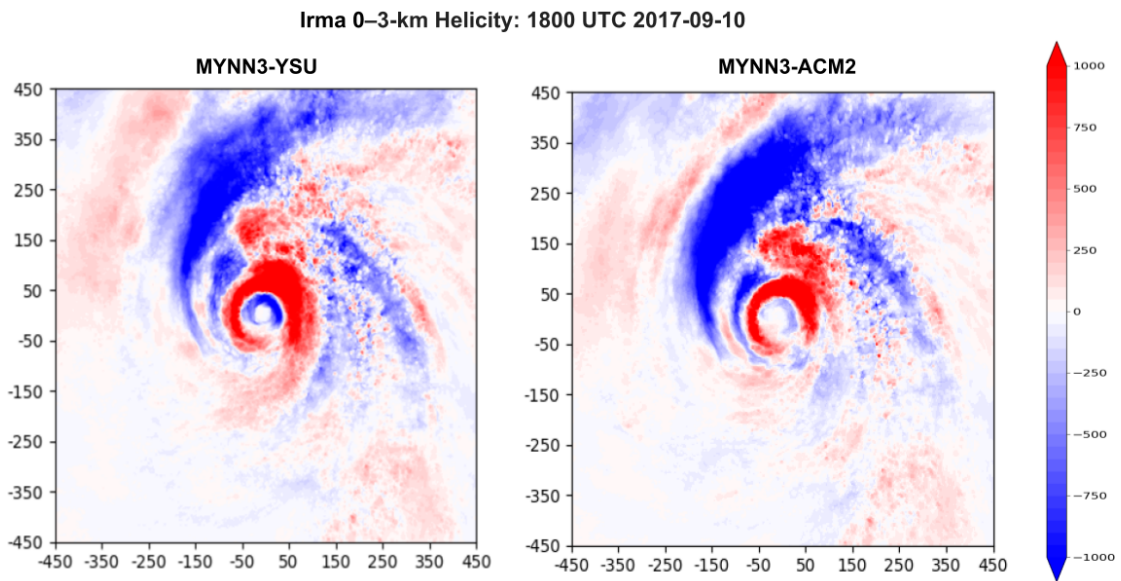




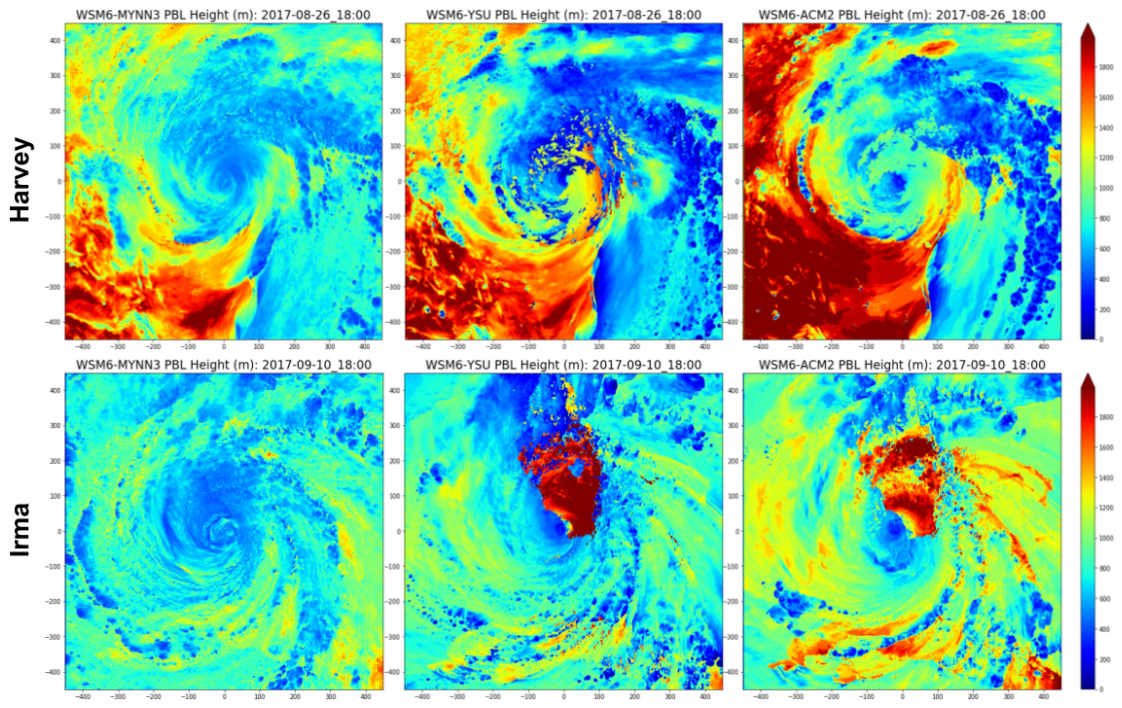
955 FIG. 13. 0–3-km vertical velocity difference ( $m s^{-1}$ , shaded) for the MYNN3 PBL scheme between the YSU  
 956 (left) and ACM2 (right) for Hurricane Irma (2017) at 1800 UTC 10 September.



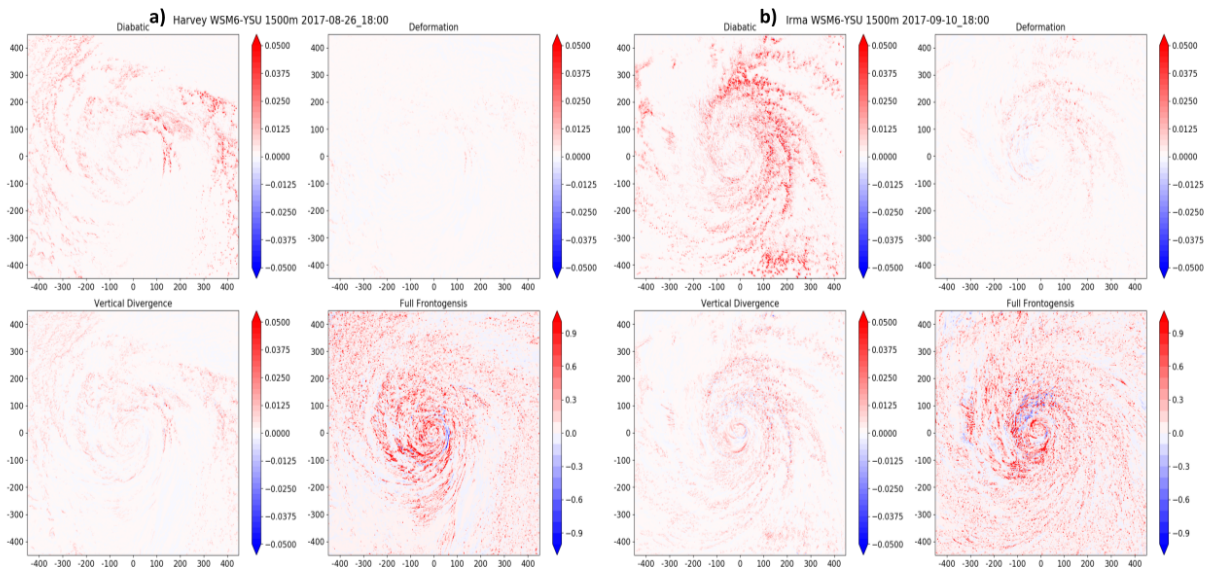
957 FIG. 14. 0–3-km helicity difference ( $m^2 s^{-2}$ , shaded) for the MYNN3 PBL scheme between the YSU (left)  
 958 and ACM2 (right) for Hurricane Harvey (2017) at 1800 UTC 26 August.



959 FIG. 15. 0–3-km helicity difference ( $m^2 s^{-2}$ , shaded) for the MYNN3 PBL scheme between the YSU (left)  
 960 and ACM2 (right) for Hurricane Irma (2017) at 1800 UTC 10 September.

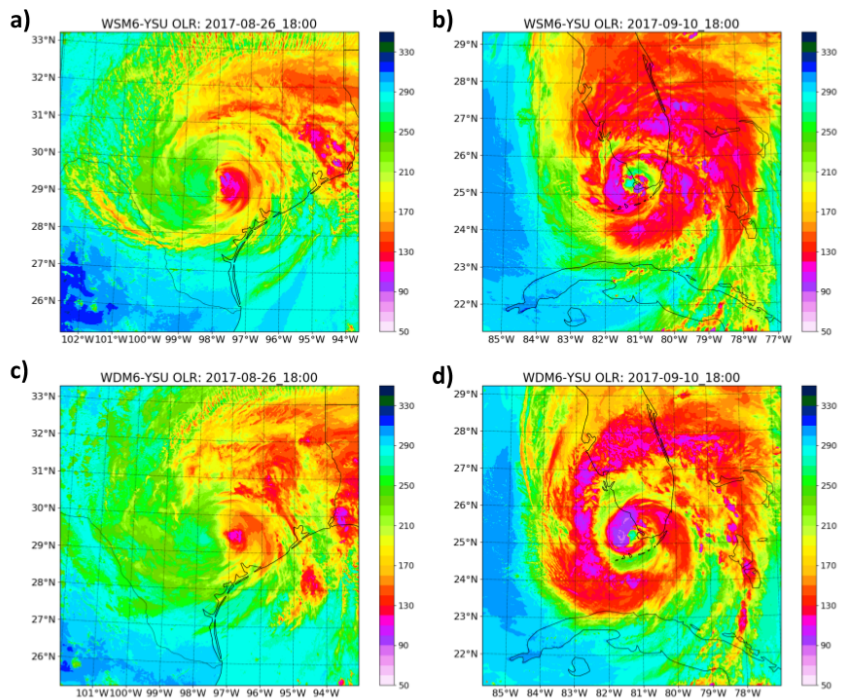


961 FIG. 16. Planetary boundary layer height ( $m$ , shaded) for the MYNN3, YSU, and ACM2 PBL schemes  
 962 in hurricanes Harvey (top) and Irma (bottom) (2017) at 1800 UTC 26 August and 1800 UTC 10 September,  
 963 respectively.

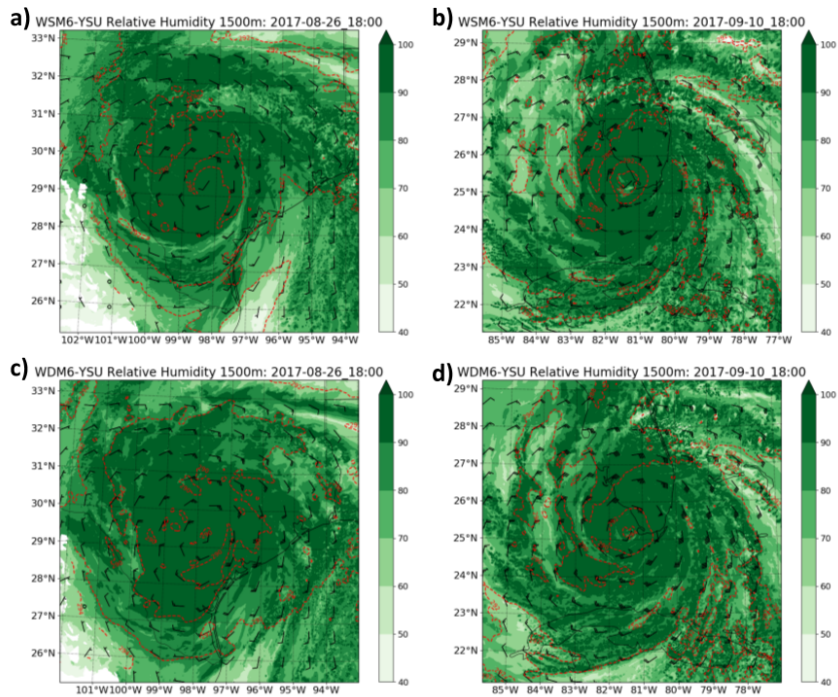


964 FIG. 17. WSM6–YSU simulation 1500-m diabatic, deformation, vertical divergence, and the full frontoge-  
 965 nesis equations ( $\frac{K}{skm}$ , shaded) at a) 1800 UTC 26 August for Hurricane Harvey (2017) and b) at 1800 UTC 10  
 966 September for Hurricane Irma (2017).





967 FIG. 18. Outgoing longwave radiation (K, shaded) for the a) WSM6 and c) WDM6 simulations for 1800 UTC  
 968 26 August for Hurricane Harvey (2017) and b) WSM6 and d) WDM6 simulations at 1800 UTC 10 September  
 969 for Hurricane Irma (2017).



970 FIG. 19. 1500-m relative humidity (% , shaded), potential temperature (K, dashed), and wind barbs ( $m s^{-1}$ ,  
 971 standard convention) for the a) WSM6 and c) WDM6 simulations for 1800 UTC 26 August for Hurricane Harvey  
 972 (2017) and b) WSM6 and d) WDM6 simulations at 1800 UTC 10 September for Hurricane Irma (2017).