The Geostrophic Wind

• On Monday, after we derived the <u>component expressions</u> for the <u>momentum equation</u>, we performed a <u>scale</u> <u>analysis on the horizontal equations</u> choosing to <u>look at</u> the <u>synoptic scale</u> of motion at 45° N, as <u>summarized</u> in the following <u>table</u>:

Table 2.1	Scale Analysis of the Horizontal Momentum Equations						
	A	В	С	D	E	F	G
x-Eq.	$\frac{Du}{Dt}$	$-2\Omega v \sin \phi$	$+2\Omega w \cos \phi$	$+\frac{uw}{a}$	$-\frac{uv \tan \phi}{a}$	$= -\frac{1}{\rho} \frac{\partial p}{\partial x}$	+ F _{rx}
y-Eq.	$\frac{Dv}{Dt}$	$+2\Omega u \sin \phi$		$+\frac{vw}{a}$	$+\frac{u^2\tan\phi}{a}$	$= -\frac{1}{\rho} \frac{\partial \boldsymbol{p}}{\partial \boldsymbol{y}}$	$\pm F_{ry}$
Scales	U^2/L	f_0U	f_0W	$\frac{UW}{a}$	$\frac{U^2}{a}$	$\frac{\delta P}{\rho L}$	$\frac{vU}{H^2}$
$(m s^{-2})$	10^{-4}	10^{-3}	10 ⁻⁶	10^{-8}	10 ⁻⁵	10^{-3}	10 ⁻¹²

• From the table, it is apparent that <u>for typical synoptic</u> <u>scale</u>, <u>midlatitude motions</u>, the <u>pressure gradient and Coriolis</u> (associated with horizontal flow) forces are of the <u>same order</u> and <u>approximately balance</u> each other; i.e.,

 $-2\Omega v \sin \phi = -\frac{1}{\rho} \frac{\partial p}{\partial x}$ in the x - direction

$$2\Omega u \sin \phi = -\frac{1}{\rho} \frac{\partial p}{\partial y}$$
 in the y-direction

- These two expressions represent geostrophic ("geo" for Earth and "strophic" for turning) balance relating the horizontal pressure and wind fields.
- Letting $f = 2\Omega \sin \phi$ and defining $\mathbf{u_g}$ and $\mathbf{v_g}$ as the zonal and meridional components of the geostrophic wind such that $\vec{V_g} = u_g \hat{i} + v_g \hat{j}$, we may write

$$-fv_g = -\frac{1}{\rho} \frac{\partial p}{\partial x} \longrightarrow v_g = \frac{1}{f\rho} \frac{\partial p}{\partial x} \text{ and } fu_g = -\frac{1}{\rho} \frac{\partial p}{\partial y} \longrightarrow u_g = -\frac{1}{f\rho} \frac{\partial p}{\partial y}$$

• We can <u>simplify</u> this <u>further</u> by <u>switching</u> from <u>height to</u> <u>geopotential height</u> coordinates. Recall $\Phi = gz$, the <u>hydrostatic equation</u> $\partial z = -\partial p/g\rho$ and the <u>height to</u> <u>pressure transform</u>:

$$-\frac{1}{\rho} \left(\frac{\partial p}{\partial x} \right)_z = -g \left(\frac{\partial z}{\partial x} \right)_p = -\left(\frac{\partial \Phi}{\partial x} \right)_p$$

we may rewrite the geostrophic wind components as

$$v_g = \frac{1}{f} \frac{\partial \Phi}{\partial x}$$
 and $u_g = -\frac{1}{f} \frac{\partial \Phi}{\partial y}$

and see the true <u>beauty of</u> the <u>geopotential height</u> in that the <u>density</u> <u>does not appear in the equations!</u>

• Thus, in vector form we may write

$$\vec{V}_g = u_g \hat{i} + v_g \hat{j} = -\frac{1}{f} \frac{\partial \Phi}{\partial y} \hat{i} + \frac{1}{f} \frac{\partial \Phi}{\partial x} \hat{j}$$

$$\vec{V}_g = \frac{1}{f} \left(-\frac{\partial \Phi}{\partial y} \hat{i} + \frac{\partial \Phi}{\partial x} \hat{j} \right)$$

$$\vec{V}_g = \frac{1}{f} \hat{k} \times \vec{\nabla}_p \Phi \longrightarrow \vec{V}_g = \frac{1}{f} \begin{bmatrix} \hat{i} & \hat{j} & \hat{k} \\ 0 & 0 & 1 \\ \frac{\partial \Phi}{\partial x} & \frac{\partial \Phi}{\partial y} & \frac{\partial \Phi}{\partial p} \end{bmatrix}$$

• As seen in the figure below, the <u>geostrophic wind blows</u> parallel to the <u>geopotential height contours</u> on a constant pressure

surface (most weather

maps) with <u>lower</u>

values of height

to the left of the 40°N

wind in the

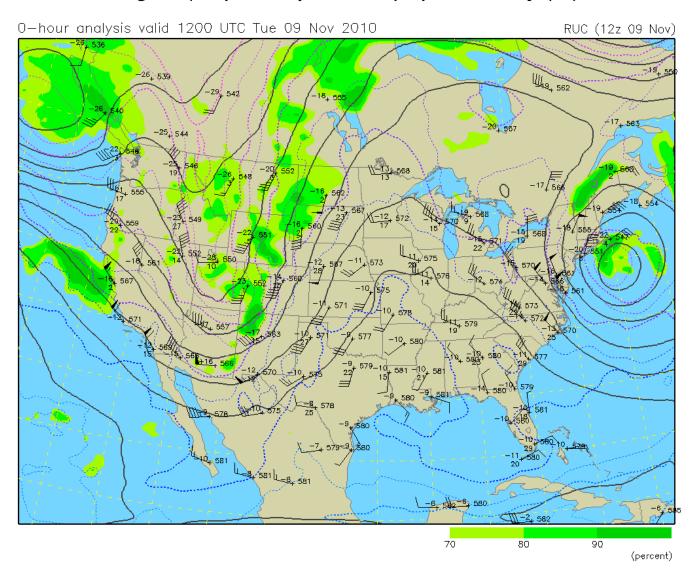
Northern Hemisphere,

and the <u>speed</u> of the <u>wind</u> is <u>inversely proportional to</u> the <u>spacing</u> of the <u>contours</u>.

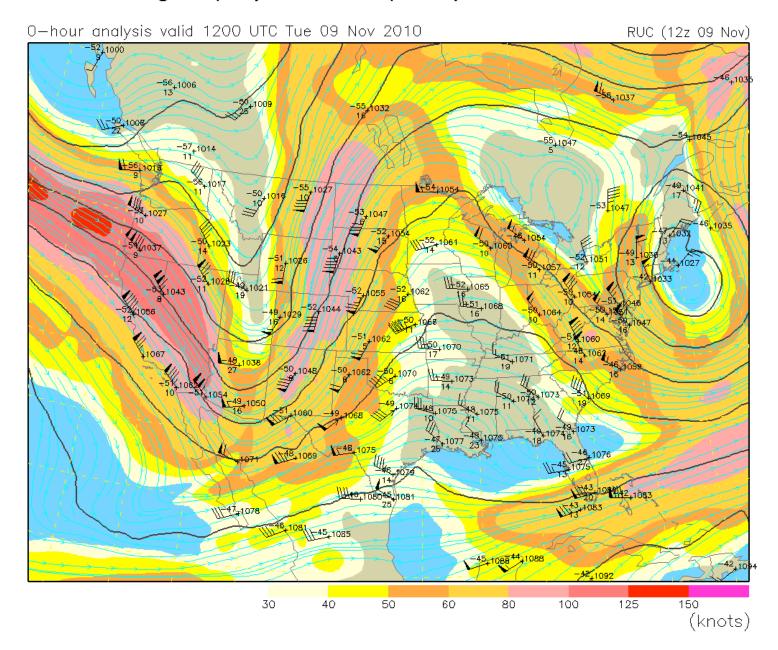
How well does this work in the atmosphere?

• Let's look at some weather maps!

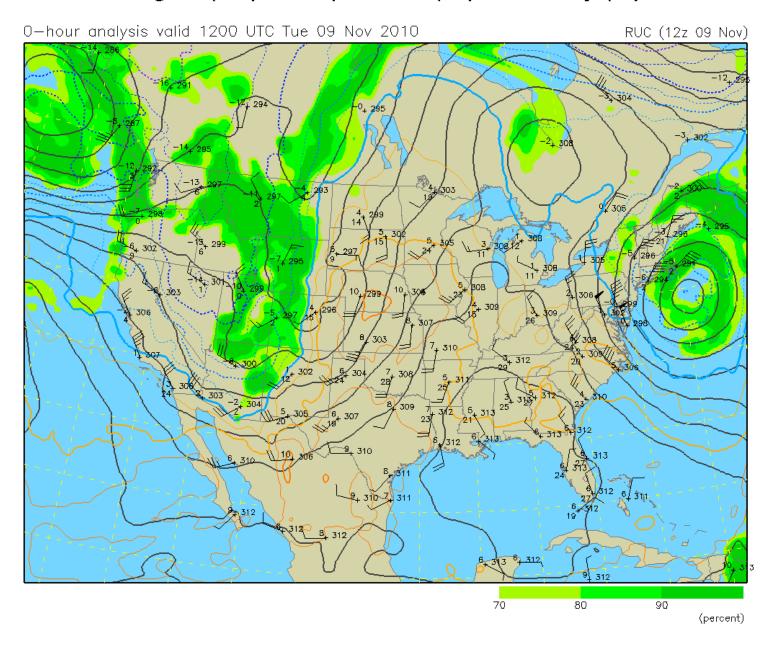
500 mb Heights (dm) / Temperature (°C) / Humidity (%)



250 mb Heights (dm) / Isotachs (knots)



700 mb Heights (dm) / Temperature (°C) / Humidity (%)



- We see that <u>geostrophic balance</u> <u>approximates the wind</u> <u>pretty darn well away from the <u>surface</u> of the <u>Earth</u> (where we have <u>friction</u>; we'll cover this in a little bit) and <u>not at the equator</u> where f = 0.</u>
- Thus, the <u>geostrophic wind</u> is <u>frequently used in place of</u> the <u>real wind</u> for <u>midlatitude</u> applications and is <u>accurate</u> for most situations to <u>within 10-15 %</u>.
- Alas, <u>making</u> the <u>assumption of geostrophy</u> does have some <u>unfortunate consequences</u> that we shall now examine.
- The <u>first consequence</u> is that there is <u>no reference to time</u> in the geostrophic wind equation; i.e., <u>the equation is diagnostic</u>, <u>not prognostic</u>.
- This means that with knowledge of the height field, we can diagnose the wind field, but we can not predict how it will evolve in time because there is no D / Dt term.

• <u>To obtain predictive equations</u> for the <u>wind field</u> we <u>must retain the</u> (D / Dt) <u>terms</u>. If we do this, the <u>resulting momentum equations</u> are

$$\frac{Du}{Dt} = fv - \frac{\partial\Phi}{\partial x}$$
 and $\frac{Dv}{Dt} = -fu - \frac{\partial\Phi}{\partial y}$

 Substituting in for the <u>height gradients</u> using the <u>definition</u> of the <u>geostrophic wind components</u> we have

$$\frac{Du}{Dt} = f(v - v_g)$$
 and $\frac{Dv}{Dt} = -f(u - u_g)$

which says that the <u>acceleration</u> of the <u>horizontal wind</u> is <u>proportional to</u> the <u>difference between the <u>actual wind</u> and the <u>geostrophic wind</u> times the Coriolis parameter.</u>

• This confirms that the <u>acceleration</u> of the <u>wind</u> is an <u>order of magnitude smaller than the Coriolis force</u> (e.g., $f(u - u_g) < f u$, with $u = 10 \text{ m s}^{-1}$ and $u_g = 9 \text{ m s}^{-1}$).

 A <u>convenient measure</u> of <u>how geostrophically</u> the <u>flow</u> is behaving is the <u>Rossby number</u>, defined as the <u>ratio of</u> the <u>acceleration</u> <u>to the Coriolis force</u>:

$$Ro = \frac{\left| \frac{D\vec{V}}{Dt} \right|}{\left| -2\vec{\Omega} \times \vec{V} \right|} = \frac{U^2/L}{f_0 U} = \frac{U}{f_0 L}$$

- Following the argument from the previous slide, when Ro is small, $V \sim V_g$ and the flow is approximately geostrophic.
- Going back to our <u>250 hPa geopotential height map</u>, where might <u>Ro not be small</u> and the flow deviating significantly from geostrophy?

- A more serious issue with using the geostrophic approximation is that there is no w equation, and if the atmosphere were indeed purely geostrophic, there would be no vertical motion.
- Let's prove it!
- Last lecture, we derived the <u>continuity equation</u>: $\vec{\nabla} \cdot \vec{V}_h = -\frac{\partial \omega}{\partial p}$
- If the atmosphere were in geostrophic balance:

$$\vec{\nabla} \cdot \vec{V}_g = -\frac{\partial \omega}{\partial p}$$

$$\frac{\partial u_g}{\partial x} + \frac{\partial v_g}{\partial y} = -\frac{\partial \omega}{\partial p}$$

Substituting in the expressions for u_g and v_g we have:

$$-\frac{\partial \omega}{\partial p} = \frac{\partial}{\partial x} \left(-\frac{1}{f_0} \frac{\partial \Phi}{\partial y} \right) + \frac{\partial}{\partial y} \left(\frac{1}{f_0} \frac{\partial \Phi}{\partial x} \right)$$

Interchanging the order of differentiation we see

$$-\frac{\partial \omega}{\partial p} = \frac{1}{f_0} \left(-\frac{\partial^2 \Phi}{\partial x \partial y} + \frac{\partial^2 \Phi}{\partial x \partial y} \right) = 0$$

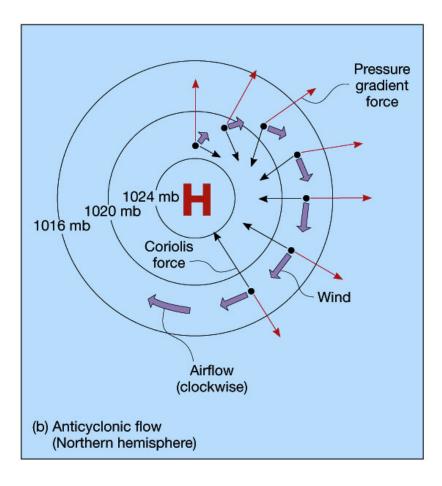
- So the geostrophic wind is non-divergent and if w = 0 at the ground and tropopause, w = 0 everywhere in between!
- Therefore, the <u>small</u> (<u>but significant</u> in some areas) <u>difference between</u> the <u>true and geostrophic winds</u> is intimately <u>tied to both</u> the <u>acceleration</u> of the horizontal <u>wind and</u> the <u>synoptic scale vertical</u> motion.

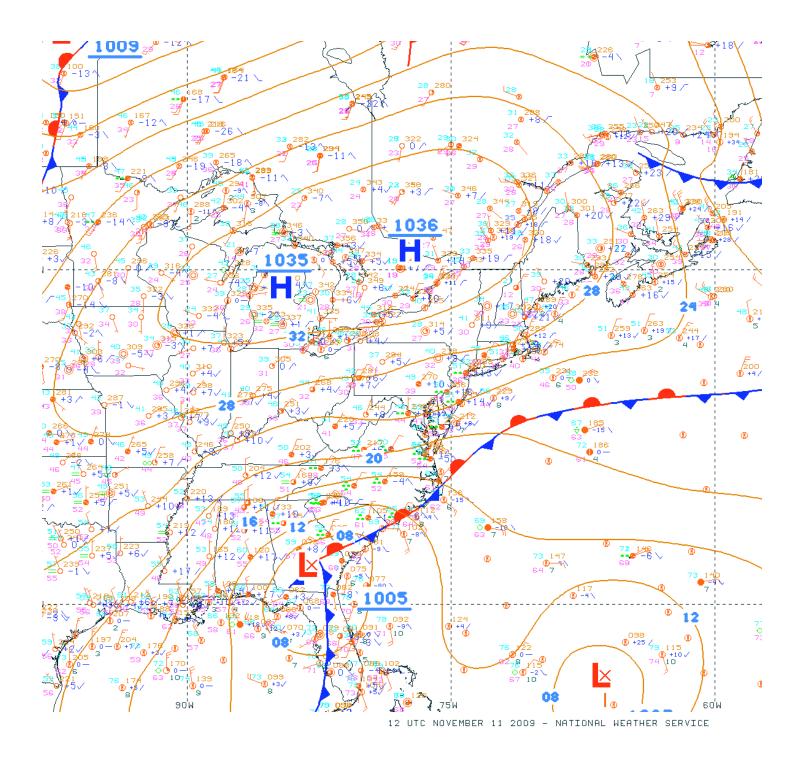
- Finally, we have seen that <u>away from jet streaks</u>, <u>the equator</u> and the <u>surface</u> of the Earth, the <u>geostrophic</u> wind is a <u>good approximation</u> for the <u>real wind</u>.
- Well, what about at the surface?

If there was no <u>friction</u>, the <u>flow</u> around a <u>high pressure</u>

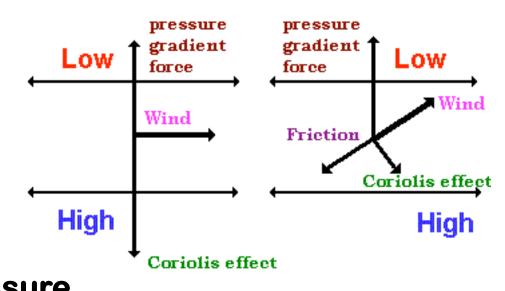
system would <u>look like</u> the <u>figure to the right</u>.

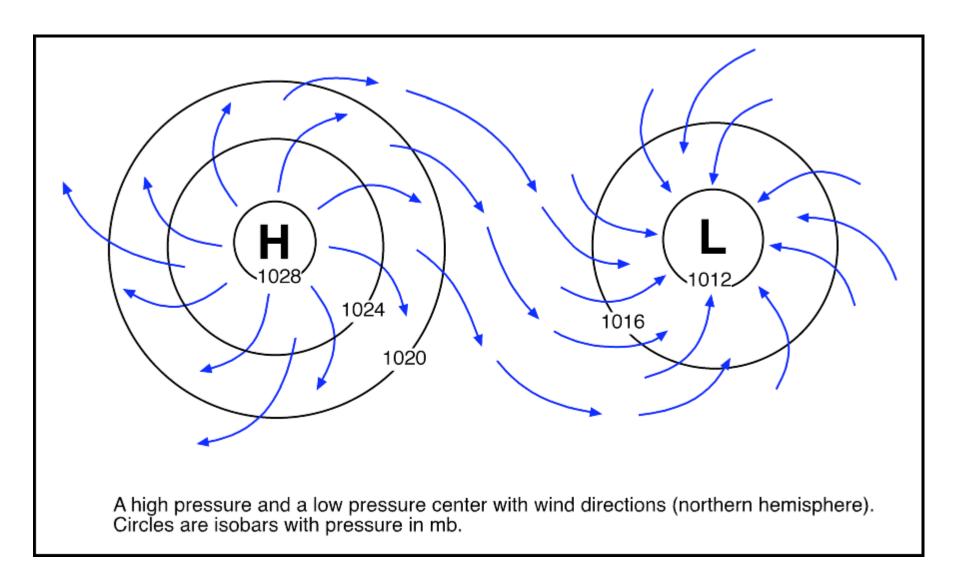
- Air would flow in a clockwise direction around high pressure systems and counterclockwise around low pressure systems in the Northern Hemisphere.
- The <u>real weather map</u> on the next slide reveals a <u>slightly</u> <u>different picture</u>.





- Examining the surface weather map, we see that <u>instead</u> of blowing parallel to the <u>isobars</u>, the <u>real surface winds</u> cross the <u>isobars</u> directed <u>out from</u> the area of <u>high</u> pressure and <u>in towards</u> the area of <u>low pressure</u>.
- This <u>cross isobaric flow</u> is the <u>result of</u> the <u>frictional force</u> <u>slowing</u> down <u>the wind</u>, making the <u>flow subgeostrophic</u>.
- This <u>reduces</u> the <u>strength of</u> the <u>Coriolis</u> force (which is directly <u>proportional to</u> the <u>wind speed</u>).
- Thus, the Coriolis force
 isn't quite strong enough
 to balance the pressure
 gradient force and the
 wind crosses the isobars
 in the direction of the
 pressure gradient force
 from higher to lower pressure.





~ Cross isobaric flow ~

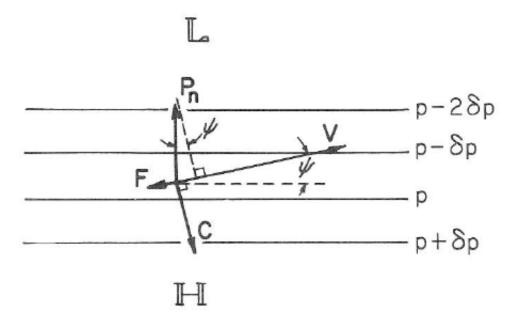
• <u>Mathematically</u>, <u>ignoring</u> the <u>acceleration</u> (DV / Dt = 0), we have a <u>three way balance</u> at the <u>surface between</u> the <u>pressure gradient</u>, <u>Coriolis</u> and <u>frictional forces</u>:

$$0 = -\frac{1}{\rho} \vec{\nabla} p - 2\vec{\Omega} \times \vec{V} + \nu \vec{\nabla}^2 \vec{V}$$

$$\frac{1}{\rho}\vec{\nabla}p = -2\vec{\Omega} \times \vec{V} + \nu \vec{\nabla}^2 \vec{V}$$

 Written this way, we can see that the <u>pressure gradient</u> force (directed

perpendicular to the isobars), must balance both the Coriolis (perpendicular to the wind), and frictional (directed opposite to the velocity) forces, as indicated on the figure.



- The <u>wind speed</u> is <u>determined by</u> the <u>requirement</u> that the <u>Coriolis force</u> be <u>just large enough to balance</u> the <u>component of</u> the <u>pressure gradient</u> force <u>in</u> the <u>direction perpendicular to the velocity</u>.
- The <u>angle between</u> V and V_g (ψ on the figure) is determined by the <u>requirement</u> that the <u>component of</u> the <u>pressure gradient force</u> in the <u>direction of</u> the <u>velocity</u> be <u>equal and opposite</u> to the <u>frictional force</u>.

