

The Vorticity Equation

①

Derivation: begin with horizontal eqn of motion:

$$\frac{d\vec{v}}{dt} = -2\vec{\omega} \times \vec{v} - \frac{1}{\rho} \nabla \rho + \vec{F}$$

recall that relative vorticity, ζ is the \hat{k} component of the curl of the 3D wind:

$$\hat{k} \cdot \vec{\omega} \equiv \hat{k} \cdot \vec{\nabla} \times \vec{v} = \frac{\partial v}{\partial x} - \frac{\partial u}{\partial y}$$

so, take $\hat{k} \cdot \vec{\nabla} \times \left(\frac{d\vec{v}}{dt} \right)$ to get $\frac{d\zeta}{dt}$ eqn

$$\textcircled{1} \hat{k} \cdot \vec{\nabla} \times \frac{d\vec{v}}{dt} = \hat{k} \cdot \vec{\nabla} \times (-2\vec{\omega} \times \vec{v}) + \hat{k} \cdot \vec{\nabla} \times \left(-\frac{1}{\rho} \nabla \rho \right) + \hat{k} \cdot \vec{\nabla} \times \vec{F}$$

expand LHS: $\hat{k} \cdot \vec{\nabla} \times \frac{d\vec{v}}{dt} \equiv \frac{\partial}{\partial x} \left(\frac{dv}{dt} \right) - \frac{\partial}{\partial y} \left(\frac{du}{dt} \right)$

$$\textcircled{2} \hat{k} \cdot \vec{\nabla} \times \frac{d\vec{v}}{dt} \equiv \frac{\partial}{\partial x} \left(\frac{\partial v}{\partial t} + u \frac{\partial v}{\partial x} + v \frac{\partial v}{\partial y} + w \frac{\partial v}{\partial z} \right) - \frac{\partial}{\partial y} \left(\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} + w \frac{\partial u}{\partial z} \right)$$

rearranging, get

$$\textcircled{3} \hat{k} \cdot \vec{\nabla} \times \frac{d\vec{v}}{dt} \equiv \frac{\partial}{\partial t} \left(\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} \right) + u \frac{\partial}{\partial x} \left(\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} \right) + v \frac{\partial}{\partial y} \left(\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} \right) + w \frac{\partial}{\partial z} \left(\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} \right) + \left(\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} \right) \left(\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} \right) + \left(\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} \right) \left(\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} \right) + \left(\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} \right) \left(\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} \right)$$

eqn can be written as

$$\textcircled{4} \hat{k} \cdot \vec{\nabla} \times \frac{d\vec{v}}{dt} = \frac{\partial \psi}{\partial t} + \vec{v} \cdot \nabla \psi + (\nabla \cdot \vec{v}) \psi + \hat{k} \cdot (\nabla \omega \times \frac{\partial \vec{v}}{\partial z})$$

(HW: fill in steps between 2 and 4)

now, deal with terms on RHS of 1

$$\hat{k} \cdot \vec{\nabla} \times \frac{d\vec{v}}{dt} = \underbrace{\hat{k} \cdot \vec{\nabla} \times (-2\vec{\Omega} \times \vec{v})}_A + \underbrace{\hat{k} \cdot \vec{\nabla} \times (-\frac{1}{\rho} \nabla p)}_B + \underbrace{\hat{k} \cdot \vec{\nabla} \times \vec{F}}_C$$

① $\hat{k} \cdot \vec{\nabla} \times (-2\vec{\Omega} \times \vec{v})$
coriolis force

$$-2\vec{\Omega} \times \vec{v} \approx 2\Omega (v \sin \phi \hat{i} - u \sin \phi \hat{j} + u \cos \phi \hat{k})$$

(neglecting ω contribution to coriolis force)

$$\hat{k} \cdot \vec{\nabla} \times (-2\vec{\Omega} \times \vec{v}) = 2\Omega \begin{vmatrix} 0 & 0 & \hat{k} \\ \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ v \sin \phi & -u \sin \phi & u \cos \phi \end{vmatrix}$$

$$\approx 2\Omega \left[\frac{\partial}{\partial x} (-u \sin \phi) - \frac{\partial}{\partial y} (v \sin \phi) \right]$$

$$= -2\Omega \frac{\partial u}{\partial x} \sin \phi - 2\Omega u \cos \phi \frac{\partial \phi}{\partial x} - 2\Omega \frac{\partial v}{\partial y} \sin \phi - 2\Omega v \cos \phi \frac{\partial \phi}{\partial y}$$

③

$$-2\Omega \frac{\partial u}{\partial x} \sin \phi \rightarrow -f \frac{\partial u}{\partial x} \text{ (recall that } f = 2\Omega \sin \phi \text{)}$$

$$-2\Omega u \cos \phi \frac{\partial \phi}{\partial x} \rightarrow 0$$

$$-2\Omega \frac{\partial v}{\partial y} \sin \phi \rightarrow -f \frac{\partial v}{\partial y}$$

$$-2\Omega v \cos \phi \frac{\partial \phi}{\partial y} \rightarrow -2\Omega v \cos \phi \frac{\partial \phi}{a \partial \phi} \text{ (} \partial y = a \partial \phi \text{, where } a \text{ is radius of earth)}$$

$$\rightarrow -v \underbrace{\frac{2\Omega \cos \phi}{a}}_{\partial f / \partial y}$$

$f = f_0 + \frac{\partial f}{\partial y} (y - y_0) \rightarrow$ linear Taylor series expansion around latitude ϕ_0

$$\rightarrow f_0 = 2\Omega \sin \phi_0$$

$$\beta \equiv \frac{\partial f}{\partial y} = \frac{\partial}{\partial y} \underbrace{(2\Omega \sin \phi)}_f = 2\Omega \cos \phi \frac{\partial \phi}{\partial y} = 2\Omega \cos \phi \frac{\partial \phi}{a \partial \phi} \rightarrow \frac{2\Omega \cos \phi}{a}$$

latitudinal variation of coriolis parameter

So term A on RHS becomes

$$-f \frac{\partial u}{\partial x} - f \frac{\partial v}{\partial y} - \beta v \rightarrow -f(\vec{\nabla} \cdot \vec{v}) - \beta v$$

$$\textcircled{b} \hat{k} \cdot \nabla \times \underbrace{\left(-\frac{1}{\rho} \nabla p\right)}_{\text{PGF}}$$

$$\equiv \hat{k} \cdot \nabla \times (-\alpha \nabla p) = \begin{vmatrix} 0 & 0 & \hat{k} \\ \partial/\partial x & \partial/\partial y & \partial/\partial z \\ -\alpha \frac{\partial p}{\partial x} & -\alpha \frac{\partial p}{\partial y} & -\alpha \frac{\partial p}{\partial z} \end{vmatrix}$$

$$= \left[\frac{\partial}{\partial x} \left(-\alpha \frac{\partial p}{\partial y}\right) \hat{i} - \frac{\partial}{\partial y} \left(-\alpha \frac{\partial p}{\partial x}\right) \hat{j} \right]$$

$$= -\frac{\partial \alpha}{\partial x} \frac{\partial p}{\partial y} - \alpha \frac{\partial^2 p}{\partial x \partial y} + \frac{\partial \alpha}{\partial y} \frac{\partial p}{\partial x} + \alpha \frac{\partial^2 p}{\partial y \partial x}$$

↑
terms sum to zero (change order of differentiation)

$$= -\frac{\partial \alpha}{\partial x} \frac{\partial p}{\partial y} + \frac{\partial \alpha}{\partial y} \frac{\partial p}{\partial x}$$

$$\rightarrow \hat{k} \cdot (\nabla p \times \nabla \alpha)$$

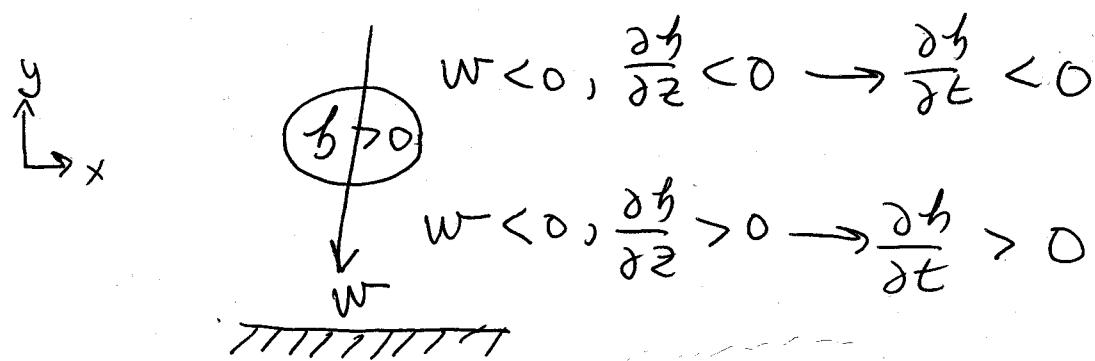
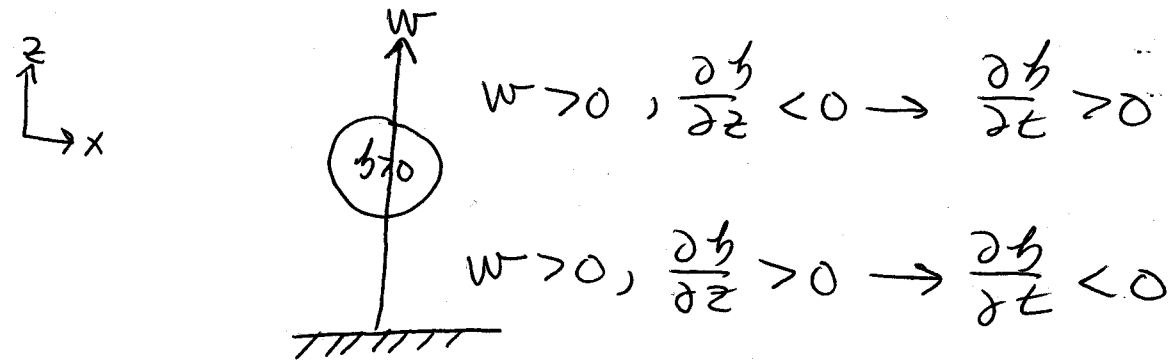
© $\hat{k} \cdot \nabla \times \vec{F} \rightarrow \hat{k}$ component of curl of friction

so all terms together:

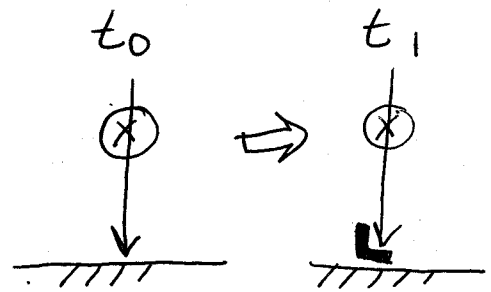
$$\frac{\partial \psi}{\partial t} + \vec{v} \cdot \nabla \psi + (\nabla \cdot \vec{v}) \psi + \hat{k} \cdot (\nabla w \times \frac{\partial \vec{v}}{\partial x}) = -f(\vec{v} \cdot \vec{v}) - \beta v + \hat{k} \cdot (\nabla w \times \frac{\partial \vec{v}}{\partial x}) + \hat{k} \cdot (\nabla p \times \nabla \alpha) + \hat{k} \cdot \nabla \times \vec{F}$$

② $-w \frac{\partial \zeta}{\partial z} > 0 \rightarrow$ cyclonic vertical vorticity advection

$-w \frac{\partial \zeta}{\partial z} < 0 \rightarrow$ anticyclonic vertical vorticity advection

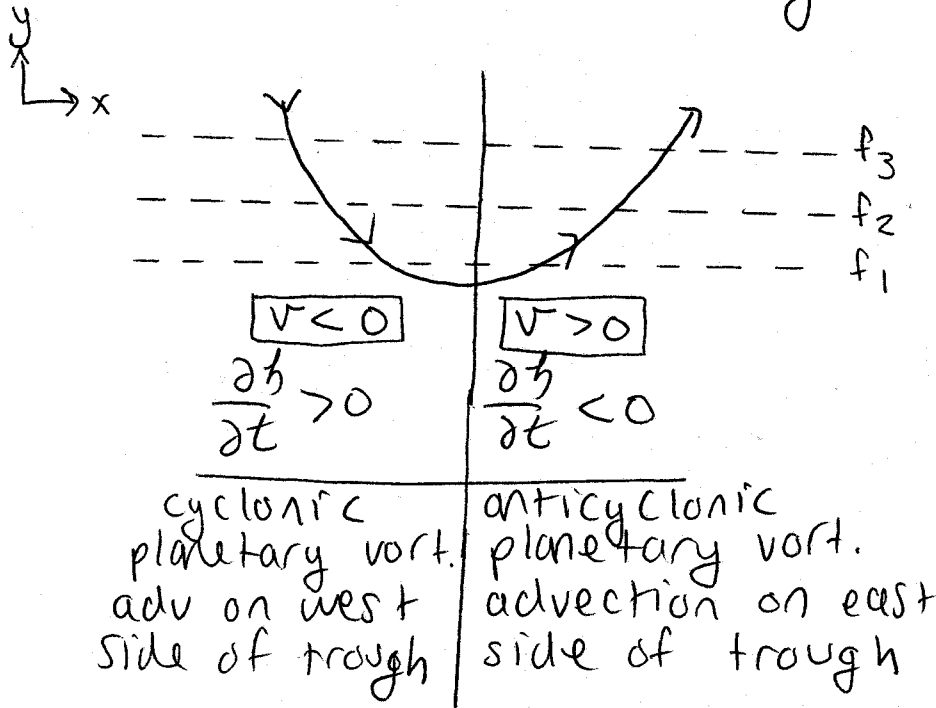


note that downdrafts ($w < 0$) can advect midlevel cyclonic vorticity down to the surface



ⓑ $\frac{\partial \zeta}{\partial t} \propto -\beta v$

$\propto -v \frac{\partial f}{\partial y} \rightarrow$ meridional advection of planetary vorticity

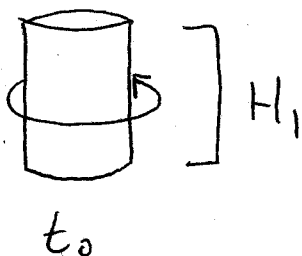


$f_3 > f_2 > f_1$
 $\frac{\partial f}{\partial y} > 0$ [f increases to N]

this term explains why a cutoff low (with little relative vorticity advection) will tend to drift westward

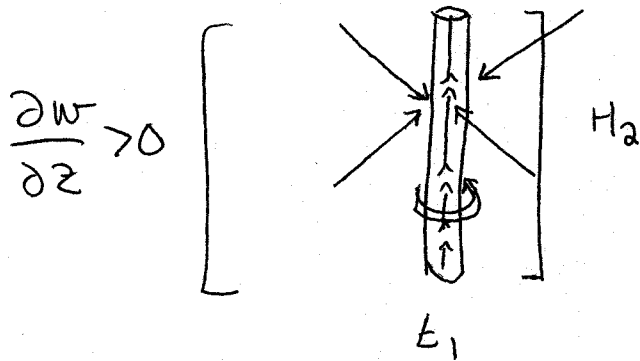
ⓒ $\frac{\partial \zeta}{\partial t} \propto -(\beta + f)(\nabla \cdot \vec{v}) \rightarrow -\beta(\nabla \cdot \vec{v}) - f(\nabla \cdot \vec{v})$
 divergence term

① consider a column with preexisting cyclonic relative vorticity



add convergence ($\nabla \cdot \vec{v} < 0$)

⑧



$H_1 < H_2$
 → column spins faster since it is stretched

fluid analog to conservation of angular momentum in rigid-body mechanics
 (ice skater pulls in arms to spin faster)

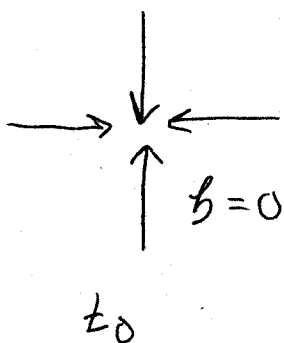
remember that $-\nabla \cdot \vec{v} = \frac{\partial \omega}{\partial z}$ for incompressible fluid
 (continuity equation / conservation of mass)

CON: $\nabla \cdot \vec{v} < 0 \rightarrow \frac{\partial \omega}{\partial z} > 0 \rightarrow$ stretching

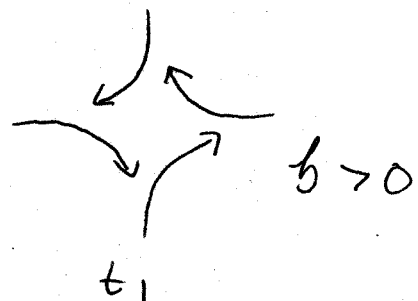
DIV: $\nabla \cdot \vec{v} > 0 \rightarrow \frac{\partial \omega}{\partial z} < 0 \rightarrow$ shrinking

② $-f(\nabla \cdot \vec{v})$

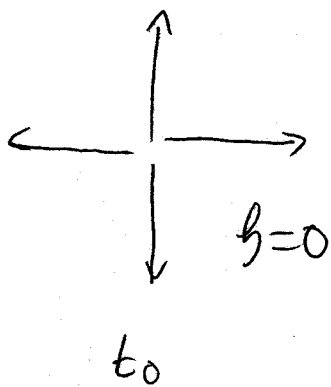
convergent scenario



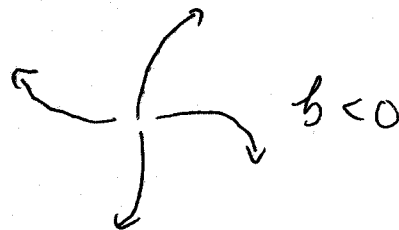
Coriolis force means that converging air acquires cyclonic relative vorticity



divergent scenario



c.f. \rightarrow diverging air acquires anticyclonic relative vorticity ⑨



CON: $\nabla \cdot \vec{v} < 0 \rightarrow \frac{\partial \zeta}{\partial t} > 0$

DIV: $\nabla \cdot \vec{v} > 0 \rightarrow \frac{\partial \zeta}{\partial t} < 0$

① $\frac{\partial \zeta}{\partial t} \propto -\hat{k} \cdot (\nabla w \times \frac{\partial \vec{v}}{\partial z})$

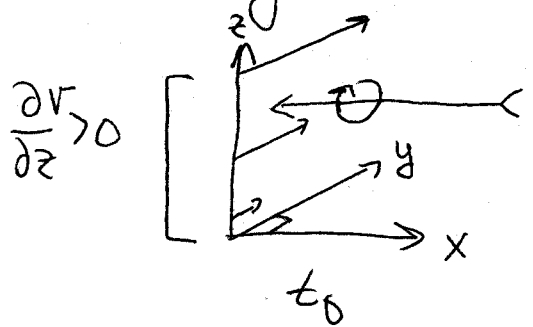
tilting term

horizontal vorticity (vertical wind shear) is tilted into the vertical by a horizontal gradient of w

$$\frac{\partial \zeta}{\partial t} \propto - \underbrace{\frac{\partial w}{\partial x} \frac{\partial v}{\partial z}}_{\text{①}} + \underbrace{\frac{\partial w}{\partial y} \frac{\partial u}{\partial z}}_{\text{②}}$$

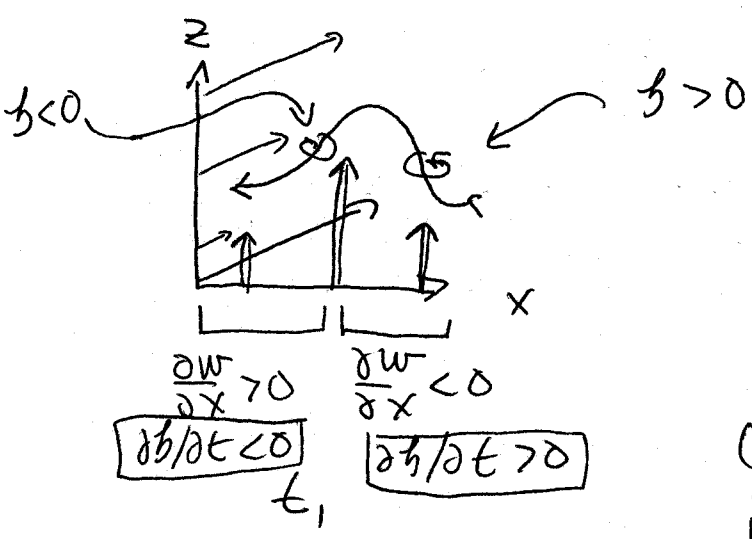
consider 1st term

have wind becoming increasingly southerly with height:



vertical wind shear creates horizontal vortex tube

add horizontal gradient of w ($\frac{\partial w}{\partial x}$);

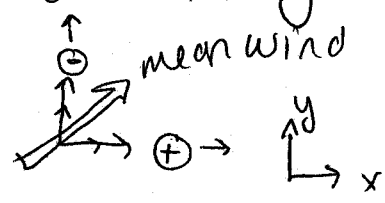


tilting of vortex tube creates cyclonic or anticyclonic relative vorticity (horizontal vorticity is tilted into vertical vorticity)

argument works the same way for 2nd term

$$\left(\frac{\partial w}{\partial y} \frac{\partial u}{\partial z} \right)$$

→ how you get supercell splitting (a cyclonic and anticyclonic rotating cell)



⊕ cyclonic supercell
⊙ anticyclonic s. cell

Ⓔ $\frac{\partial \zeta}{\partial t} \propto \hat{k} \cdot (\nabla p \times \nabla \sigma) \rightarrow$ solenoid term

when ∇p not parallel to $\nabla \sigma$, can get cyclonic or anticyclonic spin up

Ⓕ $\frac{\partial \zeta}{\partial t} \propto \hat{k} \cdot \vec{v} \times \vec{F}$

friction in PBL acts to spin down both cyclones and anticyclones through Ekman pumping

Conservation of Potential Vorticity

can simplify vorticity eqn using geostrophic assumptions and derive

$$\frac{d}{dt} \left(\frac{b_g + f}{H} \right) = 0$$

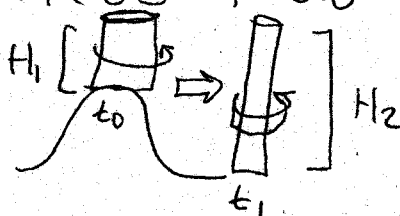
Potential vorticity

↳ absolute vorticity normalized by column height

b_g : geostrophic rel. vorticity
 f : planetary vorticity
 H : height of column

says that PV is conserved for synoptic-scale, barotropic (i.e., no horiz. temp. gradients) incompressible, frictionless flow

- if $H \uparrow$, $b_g + f$ must \uparrow
- if $H \downarrow$, $b_g + f$ must \downarrow



$H_2 > H_1$, so $(b_g + f)_2 > (b_g + f)_1$